

Interaction of an atom with electro magnetic field

Electromagnetic field is oscillating electric and magnetic fields

$$\vec{E} = \vec{E}_0 \cos(\omega t - kz)$$

$$\vec{B} = \vec{B}_0 \cos(\omega t - kz)$$

The strongest interaction of an e-m field with an electron - electro-dipole interaction:

$$U_{el} = -d \vec{E} = +e \vec{r} \cdot \vec{E}_0 \cos(\omega t - kz)$$

Important to note: the size of an atom (i.e. change of r) $\sim a_0 \sim 1 \text{ \AA}$, and for most em spectrum $1/k \sim \lambda \sim 100 \text{ nm}$ or more. Thus, we can neglect any spatial variations of e-m field across an atom: $\vec{E} = \vec{E}_0 \cos \omega t$

$$U_{el} = e \vec{r} \cdot \vec{E}_0 \cos \omega t \quad - \text{periodic perturbation}$$

Quantum version

$$\hat{H}' = e \vec{r} \cdot \vec{E}_0 \cos \omega t$$

\vec{E}_0 is a constant vector in y - z plane (since we assume that the e-m field propagates in z direction).

If we consider the first-order perturbation of this e-m field close to a transition b/w two particular levels a and b

The direction of \vec{E}_0 determine the polarization of e-m field. $\vec{E}_0 = E_0 \vec{e}_0 \quad |\vec{e}_0| = 1$

———— E_b, Ψ_b

$$\langle \Psi_a | \hat{H}' | \Psi_b \rangle = e E_0 \langle \Psi_a | \vec{r} \cdot \vec{e}_0 | \Psi_b \rangle \cos \omega t$$

$$\langle \Psi_a | \hat{H}' | \Psi_a \rangle = \langle \Psi_b | \hat{H}' | \Psi_b \rangle = 0$$

due to parity (\vec{r} is odd)

———— E_a, Ψ_a
 \hat{H}_0

$$\omega_0 = \frac{E_b - E_a}{\hbar}$$

resonance freq.

[that justifies our previous assumption about \hat{H}' being off-diagonal]

$$e \langle \Psi_a | \vec{r} \cdot \vec{e}_0 | \Psi_b \rangle = p_{ab} \quad (p \text{ is called "wiggly } p", \text{ seriously})$$

Typically p_{ab} is called a transition matrix element or transition strength, and it is determined by the spatial properties of the involved energy levels. [selection rules!]

$$\# \langle \Psi_a | \hat{H}' | \Psi_b \rangle = \frac{p_{ab} E_0 \cos \omega t}{V_{ab}}$$

Thus, we can easily relate the results from last lecture to the e-m wave excitation:

- off-resonant case, system is initially in state a

$$P_b = \left(\frac{p_{ab} E_0}{\hbar} \right)^2 \frac{\sin^2 \frac{\omega - \omega_0}{2} t}{(\omega - \omega_0)^2}$$

- resonant case $\omega = \omega_0$

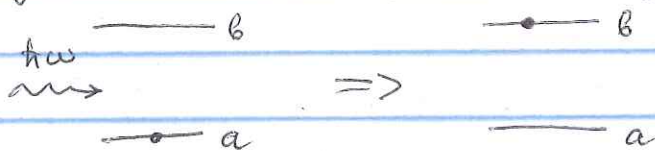
$$P_b = \sin^2 \left[\left(\frac{p_{ab} E_0}{2\hbar} \right) \cdot t \right] = \sin^2 \Omega t \quad \Omega = \frac{p_{ab} E_0}{2\hbar} \text{ Rabi freq.}$$

In general,

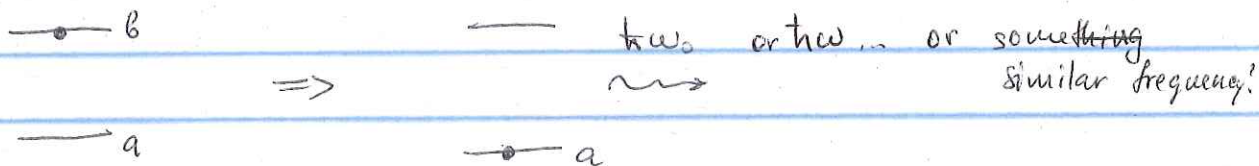
$$P_b = \left(\frac{\Omega^2}{\sqrt{\Omega^2 + (\omega - \omega_0)^2}} \right)^2 \sin^2 \left[\frac{1}{2} \sqrt{\Omega^2 + (\omega - \omega_0)^2} t \right]$$

Absorption and emission of light

It is the easiest to visualize the processes of absorption and emission if we consider light to be a stream of n particles each carrying energy $(n + \frac{1}{2}) h\nu$. Absorption of one photon gives an electron enough energy to get to a higher energy level



similarly, when an electron jumps from higher energy level to a lower energy level, an extra photon can be emitted



How to account for such energy loss/addition for a classical e-m wave? Maxwell equations!

$$\nabla \times \vec{H} = \vec{j} + \frac{\partial \vec{D}}{\partial t}$$

$$\nabla \times \vec{E} = - \frac{\partial \vec{B}}{\partial t}$$

$$\nabla \cdot \vec{D} = \rho$$

$$\nabla \cdot \vec{B} = 0$$

$$\vec{D} = \epsilon_0 \vec{E} + \vec{P}$$

$$\vec{B} = \mu_0 \vec{H} \quad \text{atomic polarization}$$

Using these we can obtain a wave equation for \vec{E} (or \vec{H})

Since $\nabla \times (\nabla \times \vec{E}) = \nabla \cdot \nabla \vec{E} - \nabla^2 \vec{E}$

$$\nabla^2 \vec{E} - \frac{1}{c^2} \frac{\partial^2 \mathcal{E}}{\partial t^2} = \mu_0 \frac{\partial^2 \vec{P}}{\partial t^2}$$

If $\vec{P} = 0$ (no atoms) — e-m wave in vacuum

atomic susceptibility
↓

usually $\vec{P} = \epsilon_0 \chi \vec{E}$
then $\frac{1}{2} \text{Im}(\chi) \cdot \frac{2\pi}{\lambda}$ — absorption coefficient

Atomic polarization $\vec{P} = \frac{\text{dipole moment}}{\text{volume}} = N \langle \vec{d} \rangle$
average dipole moment of an atom

$\langle \vec{d} \rangle = \langle \psi | -e\vec{r} | \psi \rangle$ where ψ is the wavefunction describing the state of an atom

$$\psi(t) = c_a(t) \psi_a + c_b(t) \psi_b$$

$$\langle \vec{d} \rangle = \langle c_a \psi_a + c_b \psi_b | -e\vec{r} | c_a \psi_a + c_b \psi_b \rangle = |c_a|^2 \langle \psi_a | -e\vec{r} | \psi_a \rangle + |c_b|^2 \langle \psi_b | -e\vec{r} | \psi_b \rangle + c_a c_b^* \langle \psi_b | -e\vec{r} | \psi_a \rangle + c_a^* c_b \langle \psi_a | -e\vec{r} | \psi_b \rangle$$

$$\langle \psi_a | -e\vec{r} | \psi_a \rangle = \langle \psi_b | -e\vec{r} | \psi_b \rangle = 0 \quad (\text{due to symmetry})$$

Typically $\vec{P} \parallel \vec{E}$ (and $\vec{d} \parallel \vec{E}$) so $\vec{d} = d \cdot \vec{e}_0$

$$\langle d \rangle = c_a c_b^* \langle \psi_b | -e\vec{r} \cdot \vec{e}_0 | \psi_a \rangle + c_a^* c_b \langle \psi_a | -e\vec{r} \cdot \vec{e}_0 | \psi_b \rangle =$$

$$= c_a c_b^* p_{ba} + c_a^* c_b p_{ab} = 2c_a c_b p_{ab}$$

if everything is real

Thus we can calculate the effect of atomic interaction on the amplitude of the e-m field

$$\nabla^2 \mathcal{E} - \frac{1}{c^2} \frac{\partial^2 \mathcal{E}}{\partial t^2} = \mu_0 N p_{ab} c_a c_b$$

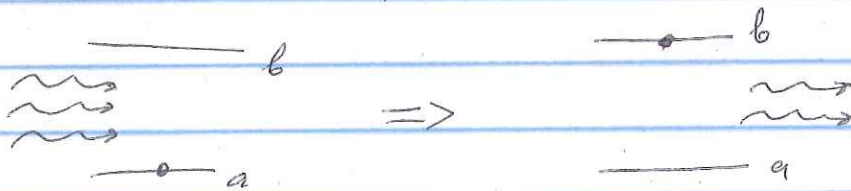
For the resonant case $c_a c_b = \cos \Omega t \sin \Omega t = \frac{1}{2} \sin 2\Omega t$

$$\nabla^2 \mathcal{E} - \frac{1}{c^2} \frac{\partial^2 \mathcal{E}}{\partial t^2} = \mu_0 N p_{ab} \sin 2\Omega t$$

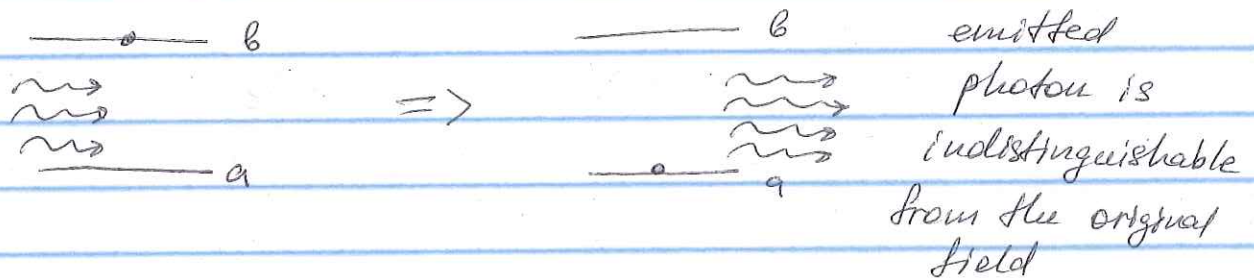
So as a function of time the amplitude is either reduced or amplified by the atomic response, but if we average the atomic polarization over time longer than $1/\Omega$, the total effect is zero. On average no energy is lost or gained!

Remember - our calculations only include what we added in. So far we only allowed one particular mode of e-m field. Thus, we only allow light to absorb and emit light that is identical to our input e-m field. Thus, the absorption and emission we consider are stimulated by the presence of the field.

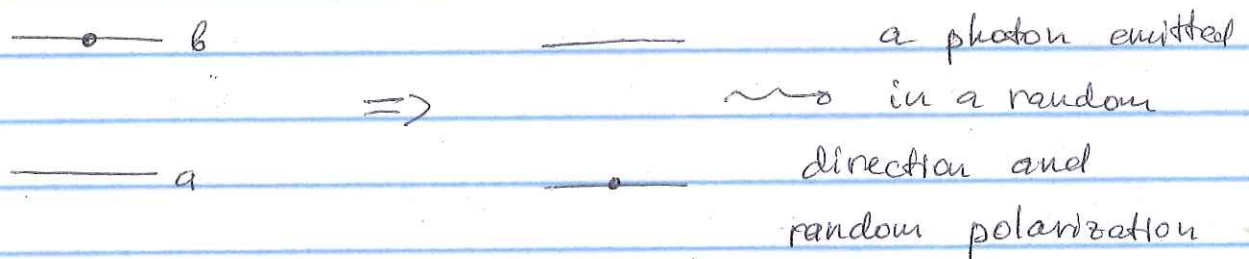
Stimulated absorption



Stimulated emission



Spontaneous emission



Energy of the emitted photon must obey energy conservation ($\approx h\nu_0$) within uncertainty relationship $\Delta E \cdot \Delta t \gtrsim \hbar$, where $\Delta E = h\nu_0$ - photon energy uncertainty, and Δt - finite lifetime of the excited state