

Homework #11

12.1 Photoelectric effect

$$E = hf - \varphi \geq 0 \quad hf_{co} = \varphi$$

$$f_{co} = \varphi/h = \frac{6.33 \text{ eV}}{4.14 \cdot 10^{-15} \text{ eV} \cdot \text{s}} = 1.53 \cdot 10^{15} \text{ Hz}$$

$$\lambda_{co} = c/f_{co} = 196 \text{ nm}$$

Such solar cell would absorb the majority of the sun spectrum

12.3

$$\text{Photon: } hf = \frac{hc}{\lambda} = E \quad \lambda = \frac{hc}{E} = \frac{1.24 \cdot 10^{-6} \text{ eV} \cdot \text{m}}{1 \text{ eV}}$$

$$= 1.24 \mu\text{m}$$

$$\text{Electron: } E^2 = p^2c^2 + m^2c^4 \quad K = E - mc^2$$

$$K \ll mc^2 \quad (1 \text{ eV} \ll 0.5 \text{ MeV}) \quad - \text{slow (non-relativistic) } \bar{e}$$

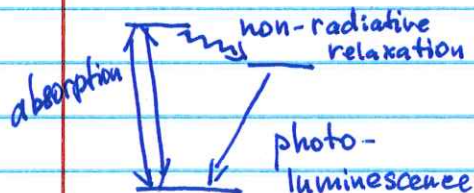
$$K \approx p^2/2m \quad p = \sqrt{2mK}$$

$$\lambda = \frac{h}{p} = \frac{h}{\sqrt{2mK}} = \frac{hc}{\sqrt{2(mc^2) \cdot K}} = \frac{1.24 \cdot 10^{-6} \text{ eV} \cdot \text{m}}{\sqrt{2 \cdot 5.10^5 \text{ eV} \cdot 1 \text{ eV}}}$$

$$\lambda = 1.23 \cdot 10^{-9} \text{ m}$$

12.12

When light is absorbed, part of photon's energy can be used for non-radiative processes, and the emitted photon has lower energy



$$h\nu_{\text{emitted}} < h\nu_{\text{abs}} \\ \lambda_{\text{emitted}} > \lambda_{\text{absorbed}}$$

A1

Assuming $m = 80 \text{ kg}$ and $v = 5 \text{ km/hr} =$

$$p = mv = 111 \text{ kg}\cdot\text{m/s}$$

$$\lambda = h/p = \frac{2\pi \cdot 10^{-34} \text{ J}\cdot\text{s}}{111 \text{ kg}\cdot\text{m/s}} \approx 5.6 \cdot 10^{-36} \text{ m}$$

To experience diffraction the wavelength of an incoming "wave" should be comparable with the distances b/w the "slits"

$$\lambda \sim 10^{-10} \text{ m} \Rightarrow p \approx h/\lambda = 6.3 \cdot 10^{-24} \text{ kg}\cdot\text{m/s}$$

$$v \approx 8 \cdot 10^{-26} \text{ m/s}$$

Much smaller than thermal velocity of atoms/molecules, so all quantum effects are smeared.

A2

$$|\psi\rangle = \frac{4}{5} |h\rangle + \frac{3}{5} |t\rangle$$

$$P_{\text{head}} = \left(\frac{4}{5}\right)^2 = 0.64$$

$$P_{\text{tail}} = \left(\frac{3}{5}\right)^2 = 0.36$$

Equal probability $|\psi_1\rangle = \frac{1}{\sqrt{2}} |h\rangle + e^{i\varphi} \frac{1}{\sqrt{2}} |t\rangle$

where φ is any real argument $[0, 2\pi]$

13.6

$$hf = E_3 - E_1 = -\frac{R_y}{3^2} + \frac{R_y}{1} = \frac{8}{9} R_y = 12.1 \text{ eV}$$

$$\frac{hc}{\lambda} \Rightarrow \lambda = \frac{hc}{E_3 - E_1} = \frac{1.24 \cdot 10^{-6} \text{ eV}\cdot\text{m}}{12.1 \text{ eV}} = 102 \text{ nm}$$