LIGO and discovery of the gravitational waves



and Eugeniy E. Mikhailov



April 13, 2018

Eugeniy Mikhailov (W&M)

LIGO and GW

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2017/10/03 Nobel prize in Physics

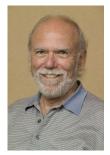
"for decisive contributions to the LIGO detector and the observation of gravitational waves"



Rainer Weiss



Kip S. Thorne

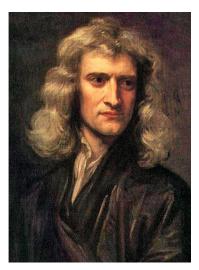


Barry C. Barish

■LIGO LIGO Scientific Collaboration



Newton's laws 1686





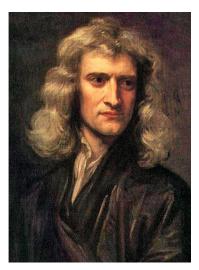
$$F_g = G rac{m_1 m_2}{r^2}$$

Laws of motion and law of gravitation solved problems of astronomy and terrestrial physics.

- eccentric orbits
- tides

• perturbation of moon orbit due to sun Unified the work of Galileo, Copernicus and Kepler.

Newton's laws 1686





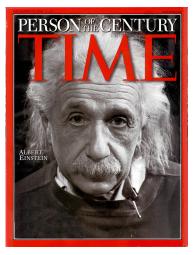
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Time is not in the formula

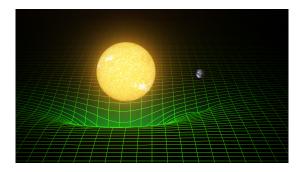


The General Theory of Relativity and theory of Gravity (1915)

- No absolute motion thus only relative motion
- Space and time are not separate thus four dimensional space-time
- Gravity is not a force acting at a distance thus warpage of space-time

General relativity

- A geometric theory connecting matter to spacetime
- Matter tells spacetime how to curve
- Spacetime tells matter how to move



important predictions

- $\bullet\,$ Light path bends in vicinity of massive object \rightarrow confirmed in 1919
- Gravitational radiation (waves) \rightarrow confirmed indirectly in 1974

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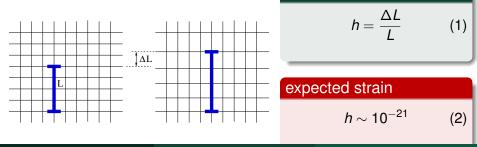
Gravitational waves (GW)

- Predicted by the General Theory of Relativity
- Generated by aspherical mass distribution
- Induce space-time ripples which propagate with speed of light



Strain - strength of GW

GW stretch and squeeze space-time thus move freely floating objects

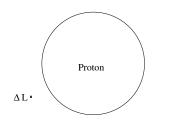


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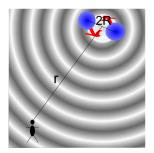
GW stretch and squeeze space-time thus move freely floating objects



Strain - strength of GW
$$h = \frac{\Delta L}{L} \qquad (1)$$
expected strain

$$h \sim 10^{-21}$$
 (2)

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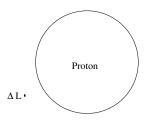


$$M_c = \frac{(m_1 m_2)^{3/5}}{(m_1 + m_2)^{1/5}}$$

$$h = 4 \frac{G}{c^2} \frac{M_c}{r} \left(\frac{G}{c^3} \pi f M_c\right)^{2/3}$$

Assuming $m_1 = m_2 = m$ and recalling that

$$f^2 \sim Gm/R^3$$



 $h \sim rac{G^2 m^2}{rRc^4} \sim rac{R_s^2}{Rr}$ Where R_s is Schwarzschild radius of the mass m $R_s = rac{2Gm}{c^2}$

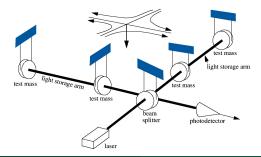
we obtain

Laser Interferometer Gravitational-wave Observatory

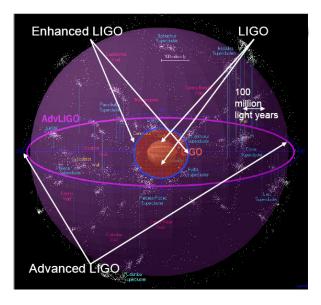






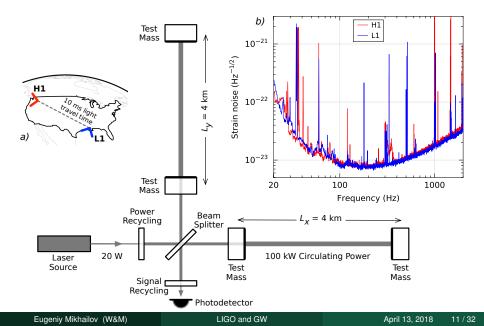


From LIGO to advanced LIGO

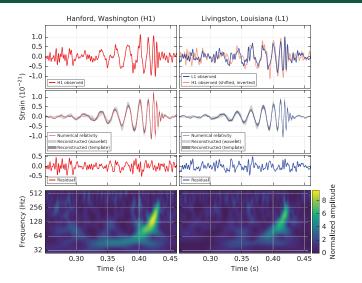


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advanced LIGO detector summary



GW signal at 09:50:45 UTC on 14 September 2015



LIGO Scientific Collaboration, "Observation of Gravitational Waves from a Binary Black Hole Merger", Phys. Rev. Lett., 116, 061102, (2016).

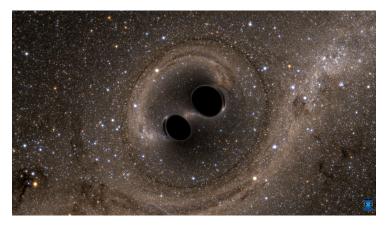
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The sound of gravitational wave and simulated sky

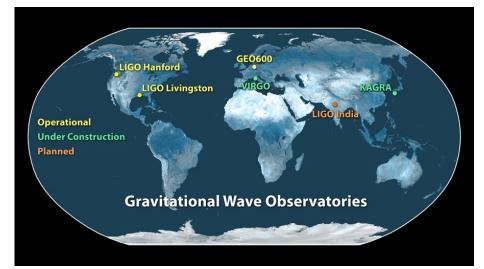
- The Sound of Two Black Holes Colliding
- Two Black Holes Merge into One



Two black holes with 29 and 36 solar masses merged about 1.3 billion years ago

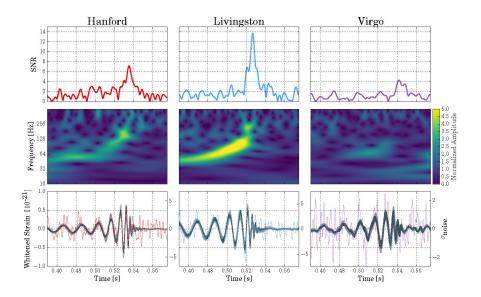
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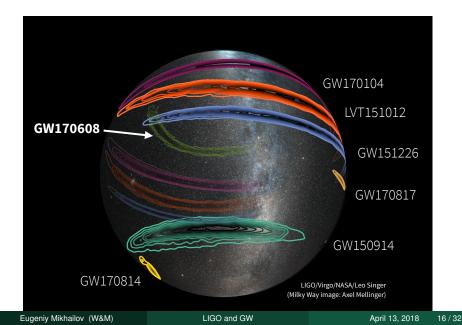
World wide network of detectors



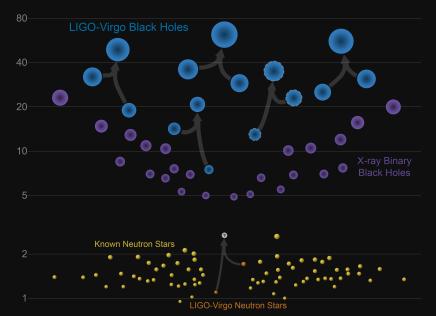
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GW170814 triple detection

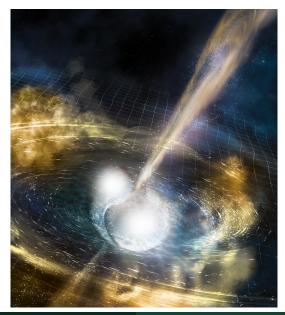




Masses in the Stellar Graveyard



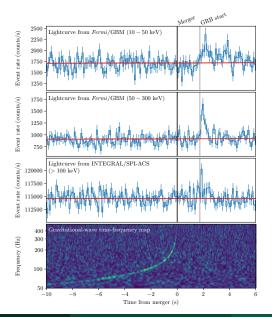
GW170817-kilonova artistic depiction



Simulation movie https://youtu. be/V6cm-0bwJ98

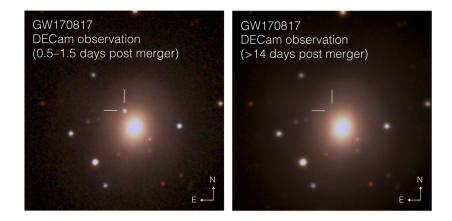
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GW170817-kilonova: two neutron stars collision



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GW170817-kilonova: two neutron stars collision



GW170817: Coherent Spectrogram and Audio https://dcc.ligo.org/LIGO-G1701924
GW170817: Fermi and LIGO signals https://wiki.ligo.

org/pub/EPO/GW170817/GBM_GW170817_small.mov

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Crab nebula, supernova 1054 remnants

		CRAB	NEBULA			
	-			1	•	
RADIO	INFRARED	VISIBLE LIGHT	ULTRAVIOLET	X-RAYS	GAMMA R	AYS
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Inside the tube



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Seismic isolation



Photo from LIGO Magazine http://www.ligo.org/magazine/

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Part of large system



Photo from LIGO Magazine http://www.ligo.org/magazine/

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Work in chamber



Inside vacuum chamber

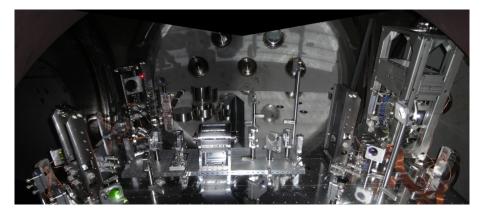


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Mode cleaner optics



Large optics suspension



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Mirror



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Inner test mass

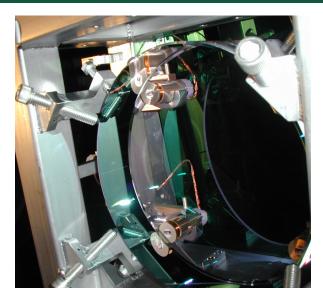


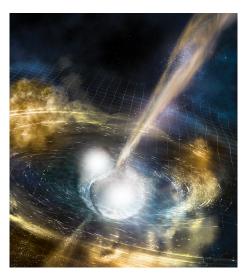
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We can detect stars collisions and ...



Summary



- In 2015 we detected the first Gravitational Wave
- Now we are talking about GW astronomy