

# LIGO and discovery of the gravitational waves



and  
Eugeniy E. Mikhailov

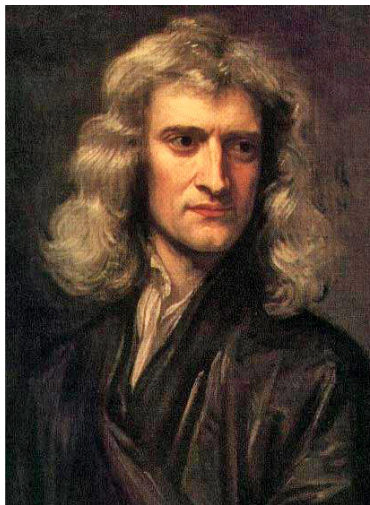


October 15th, 2016



# LIGO Scientific Collaboration





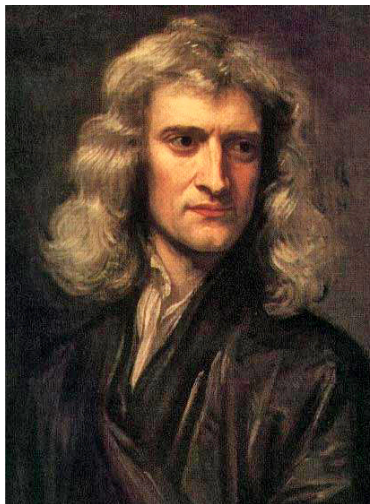
$$F_g = G \frac{m_1 m_2}{r^2}$$

Laws of motion and law of gravitation solved problems of astronomy and terrestrial physics.

- eccentric orbits
- tides
- perturbation of moon orbit due to sun

Unified the work of Galileo, Copernicus and Kepler.

# Newton's laws 1686



$$F_g = G \frac{m_1 m_2}{r^2}$$

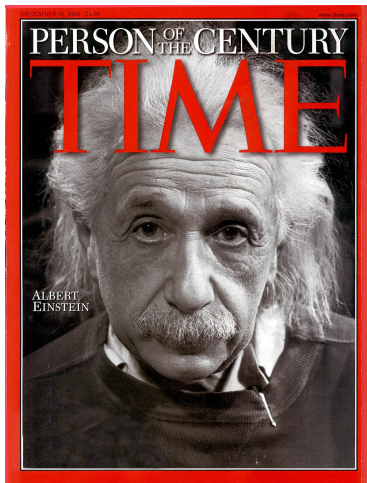
Laws of motion and law of gravitation solved problems of astronomy and terrestrial physics.

- eccentric orbits
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Unified the work of Galileo, Copernicus and Kepler.

Did not explained precession of Mercury orbit



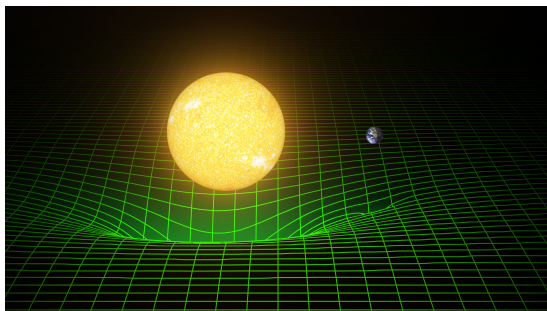


## The General Theory of Relativity and theory of Gravity (1915)

- No absolute motion  
thus only relative motion
- Space and time are not separate  
thus four dimensional space-time
- Gravity is not a force acting at a distance  
thus warpage of space-time

# General relativity

- A geometric theory connecting matter to spacetime
- Matter tells spacetime how to curve
- Spacetime tells matter how to move



## important predictions

- Light path bends in vicinity of massive object → confirmed in 1919
- Gravitational radiation (waves) → confirmed **indirectly** in 1974

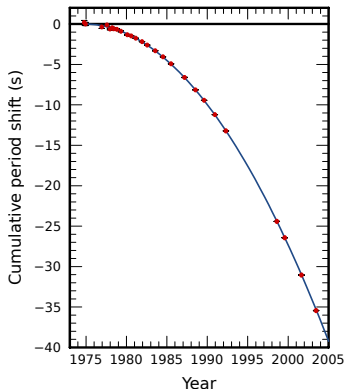
# Indirect observation of gravitational wave

Emission of gravitational radiation from pulsar PSR1913+16 leads to loss of orbital energy.

- orbital period decreased by 36 sec from 1975 to 2005
- measured to 50 ms accuracy
- deviation grows quadratically with time

This can be explained by general relativistic effects: J.H. Taylor and J.M. Weisberg, *Astrophysical Journal*, Part 1, vol. 253, Feb. 15, 1982, p. 908-920.

Nobel prize in 1993 to Hulse and Taylor



# Astrophysics with GWs vs. E&M

## E&M (photons)

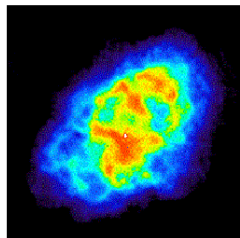
- Space as medium for field
- Accelerating charge
- Absorbed, scattered, dispersed by matter
- 10 MHz and up
- Light = not dark (but >95% of Universe is dark)

## GW

- Spacetime itself ripples
- Accelerating aspherical mass
- Very small interaction; matter is transparent
- 10 kHz and down
- Radiated by dark mass distributions

# New view to the universe

## Crab Nebula: Remnant of an Exploded Star (Supernova)



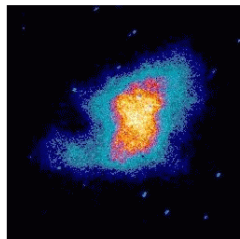
Radio wave (VLA)



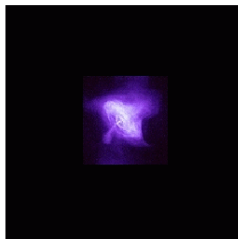
Infrared radiation (Spitzer)



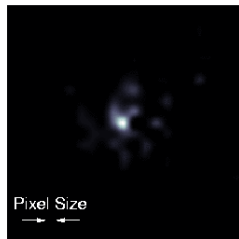
Visible light (Hubble)



Ultraviolet radiation (Astro-1)



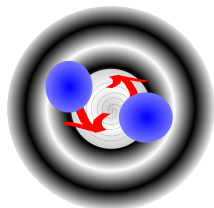
Low-energy X-ray (Chandra)



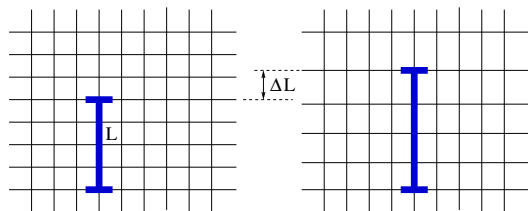
High-energy X-ray (HEFT)

# Gravitational waves (GW)

- Predicted by the General Theory of Relativity
- Generated by aspherical mass distribution
- Induce space-time ripples which propagate with speed of light



GW stretch and squeeze space-time thus move freely floating objects



Strain - strength of GW

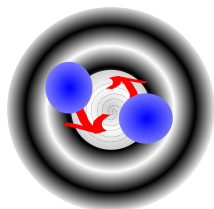
$$h = \frac{\Delta L}{L} \quad (1)$$

expected strain

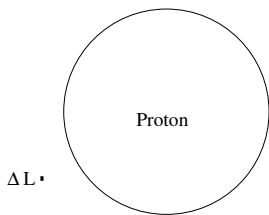
$$h \sim 10^{-21} \quad (2)$$

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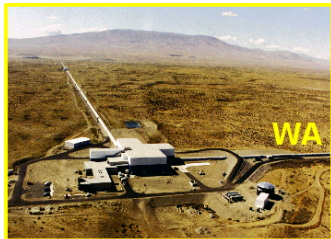
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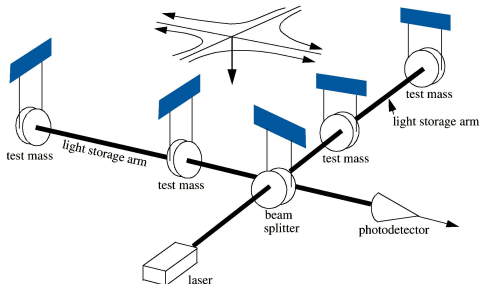
expected strain

$$h \sim 10^{-21} \quad (2)$$

# Laser Interferometer Gravitational-wave Observatory

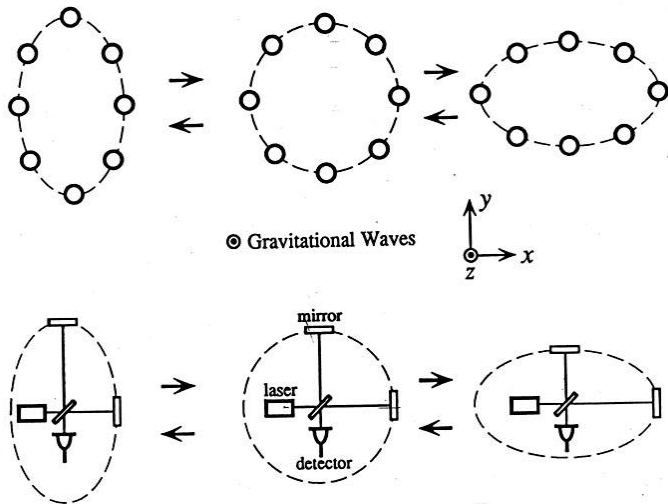


- $L = 4 \text{ km}$
- $h \sim 10^{-23}$

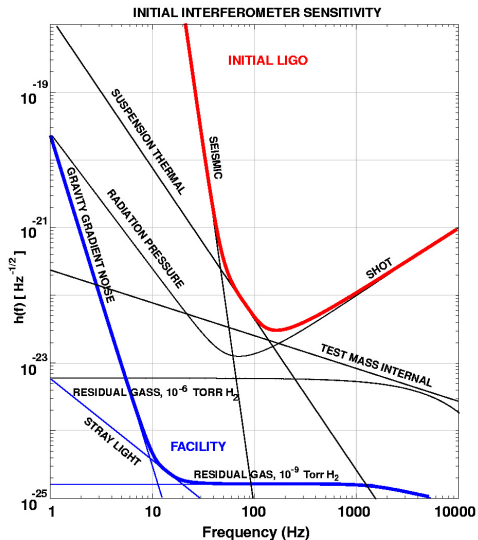




# GW acting on matter



# Initial LIGO sensitivity goal and noise budget



## Displacement noise

- seismic
- thermal suspension
- thermal Brownian
- radiation pressure noise

## Detection noise

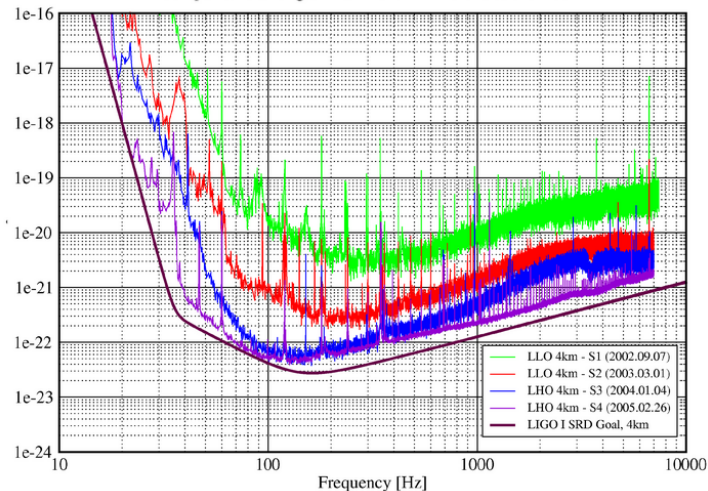
- electronics
- shot noise

# LIGO sensitivity, S1-S4 runs

## Best Strain Sensivities for the LIGO Interferometers

Comparisons among S1 - S4 Runs

LIGO-G050482-00-Z

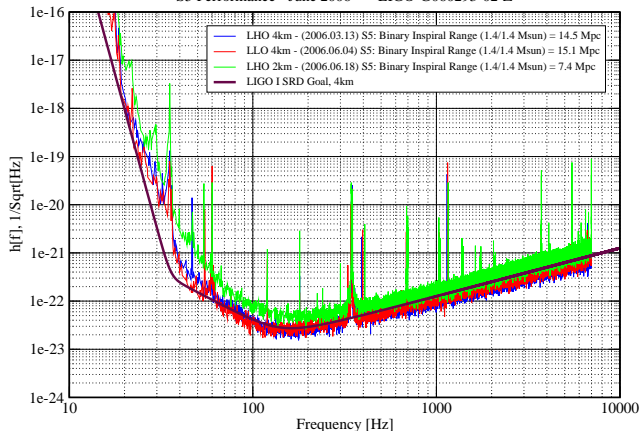


Inspiral search range during S4 was 8Mpc

# LIGO sensitivity, S5 run, June 2006

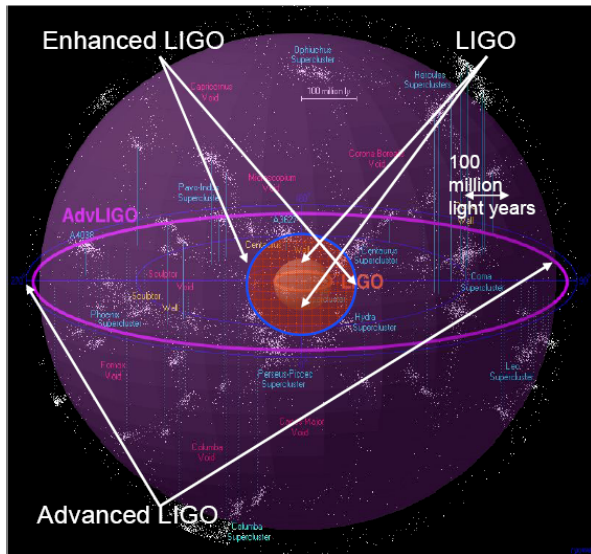
## Strain Sensitivity for the LIGO Interferometers

S5 Performance - June 2006 LIGO-G060293-02-Z

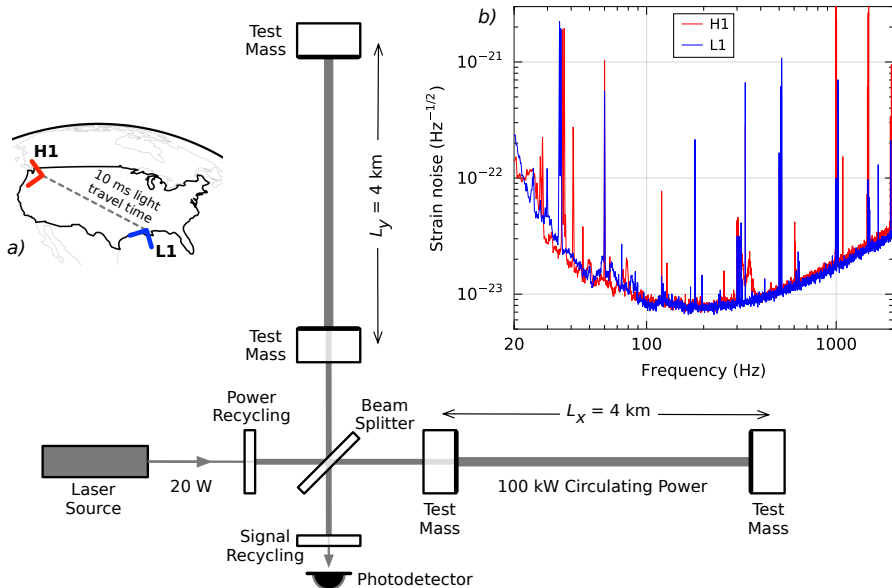


Inspiral search range during S5 is 14Mpc

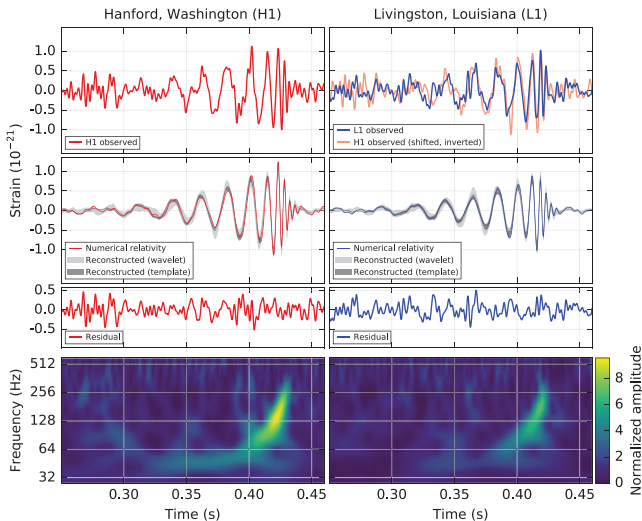
# From LIGO to advanced LIGO



# advanced LIGO detector summary



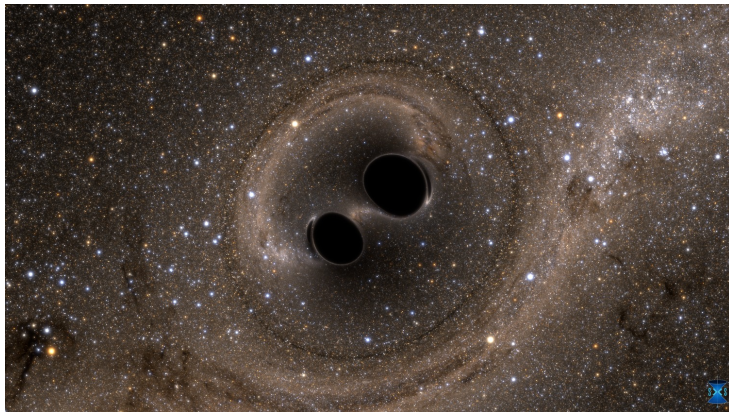
# GW signal at 09:50:45 UTC on 14 September 2015



LIGO Scientific Collaboration, "Observation of Gravitational Waves from a Binary Black Hole Merger", Phys. Rev. Lett., 116, 061102, (2016).

# The sound of gravitational wave and simulated sky

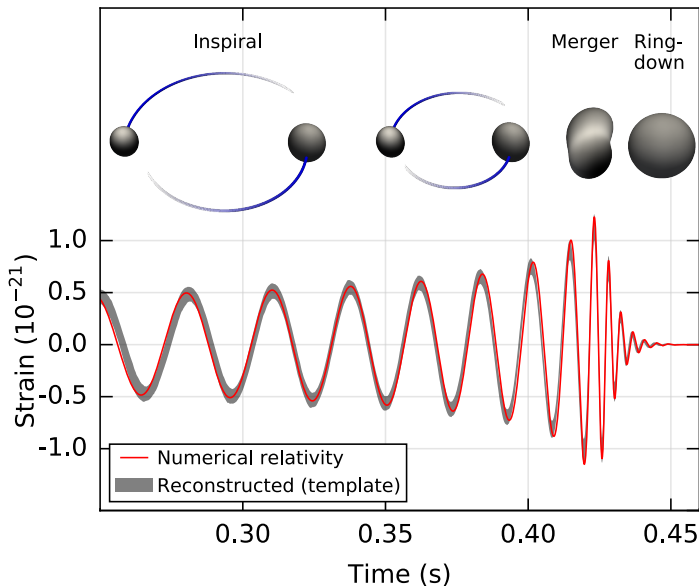
- The Sound of Two Black Holes Colliding
- Two Black Holes Merge into One



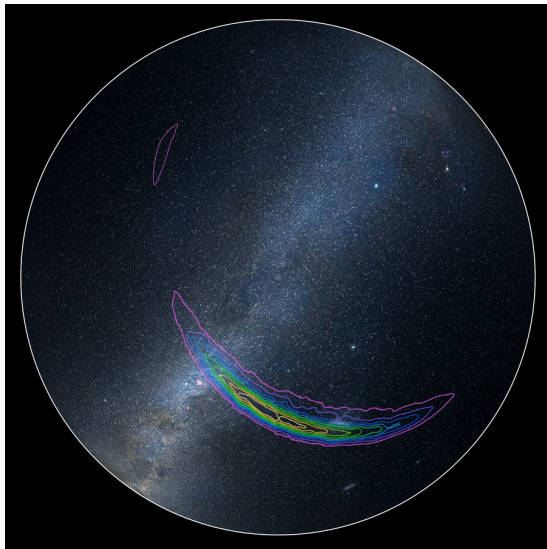
Two black holes with 29 and 36 solar masses merged about 1.3 billion years ago



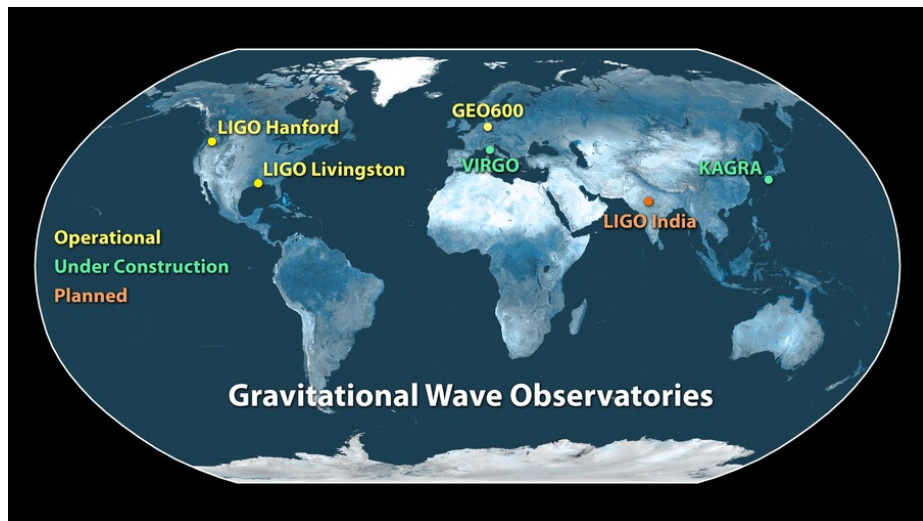
# Reconstructed signal



# GW source location at the southern hemisphere sky



# World wide network of detectors



# Confirmed GW detections

Event name	GW150914	GW151226
Mass 1	$36 M_{\odot}$	$14.2 M_{\odot}$
Mass 2	$29 M_{\odot}$	$7.5 M_{\odot}$
Final mass	$62 M_{\odot}$	$20.8 M_{\odot}$

LIGO Scientific Collaboration:

"Observation of Gravitational Waves from a Binary Black Hole Merger", Phys. Rev. Lett., 116, 061102, (2016).

"GW151226: Observation of Gravitational Waves from a 22-Solar-Mass Binary Black Hole Coalescence", Phys. Rev. Lett., 116, 241103, (2016).

# Seismic isolation



Photo from LIGO Magazine <http://www.ligo.org/magazine/>

# Part of large system



Photo from LIGO Magazine <http://www.ligo.org/magazine/>

# Work in chamber



# Inside vacuum chamber

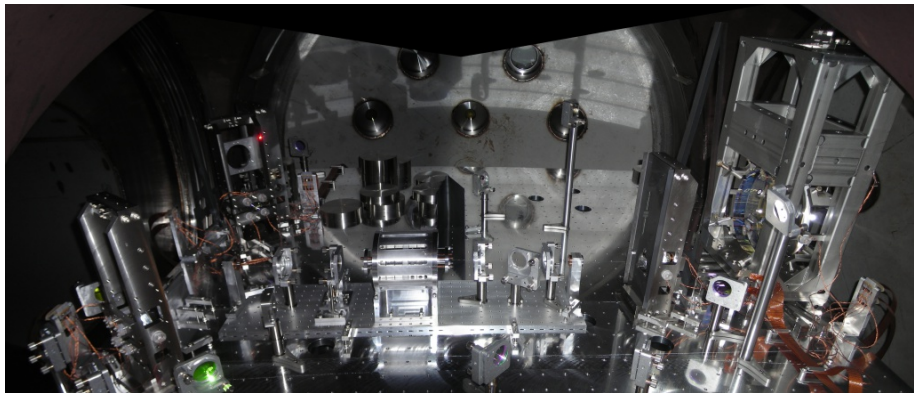


Photo from LIGO Magazine <http://www.ligo.org/magazine/>



# Mirror

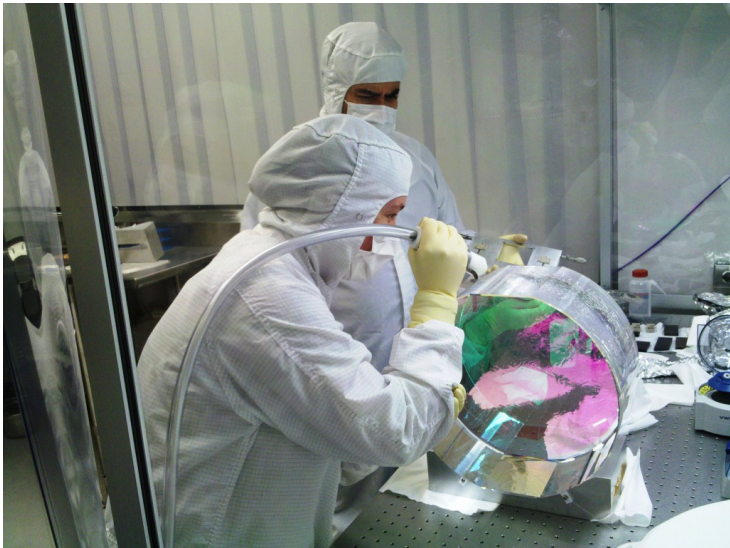


Photo from LIGO Magazine <http://www.ligo.org/magazine/>

# Inner test mass

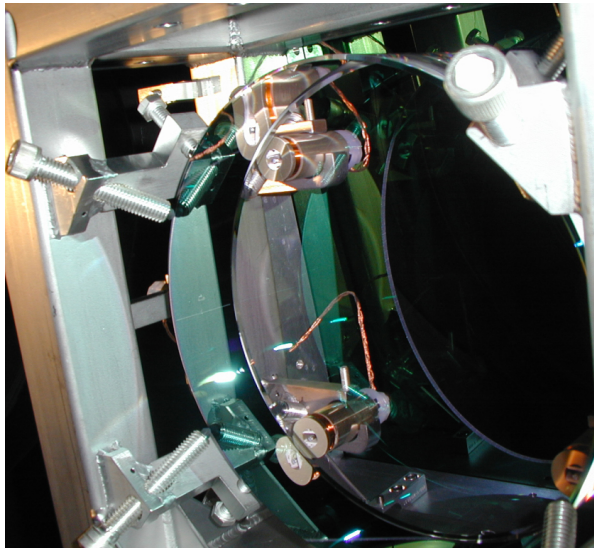


Photo from LIGO Magazine <http://www.ligo.org/magazine/>

# We can detect stars collisions and ...



- Gravitational waves exist and they are detected
- Moreover we can learn from them and do GW astronomy
- The future is in quantum noise suppression