# Squeezed states of light - generation and applications

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## From ray optics to semiclassical optics

#### Classical/Geometrical optics

- light is a ray
- which propagates straight
- cannot explain diffraction and interference



## From ray optics to semiclassical optics

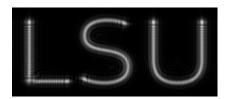
#### Classical/Geometrical optics

- light is a ray
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#### Semiclassical optics

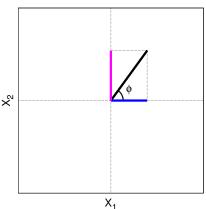
- light is a wave
- color (wavelength/frequency) is important
- amplitude (a) and phase are important,
   E(t) = ae<sup>i(kz-ωt)</sup>
- cannot explain residual measurements noise



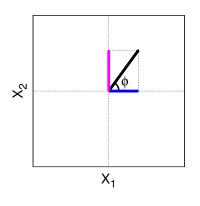


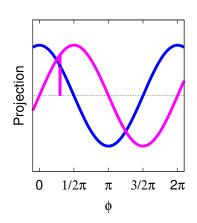
$$E(\phi) = |a|e^{-i\phi} = |a|\cos(\phi) + i|a|\sin(\phi) = X_1 + iX_2, \quad \phi = \omega t - kz$$

Detectors sense the real part of the field  $(X_1)$  but there is a way to see  $X_2$  as well

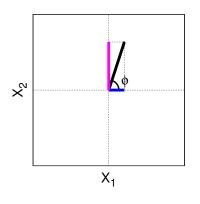


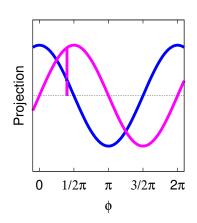
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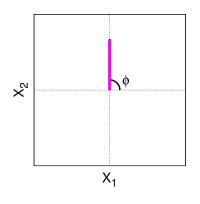


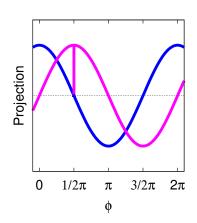
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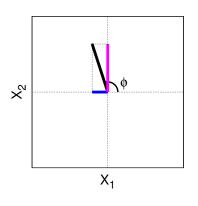


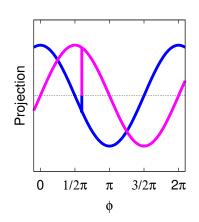
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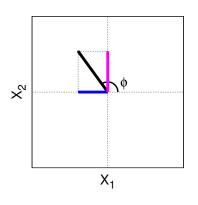


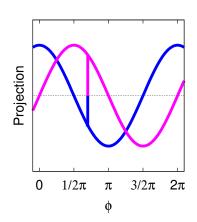
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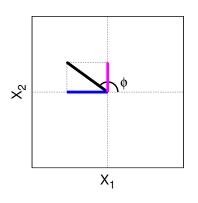


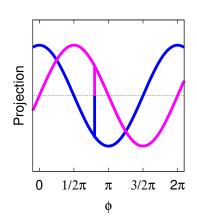
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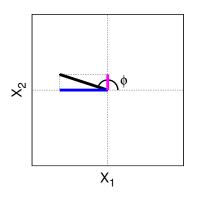


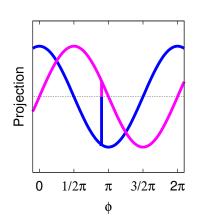
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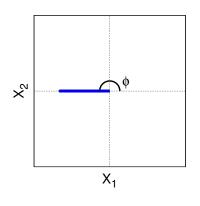


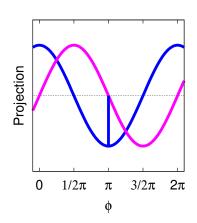
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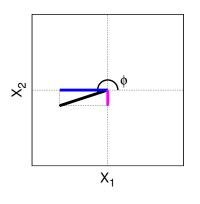


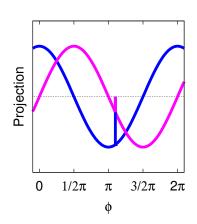
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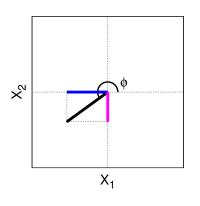


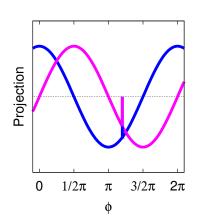
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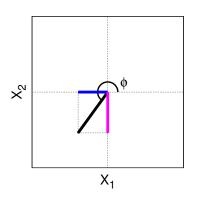


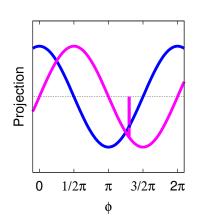
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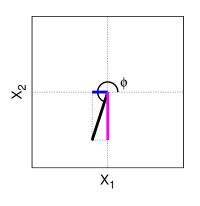


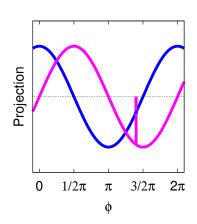
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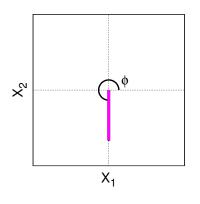


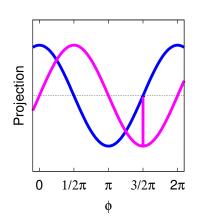
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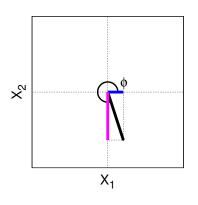


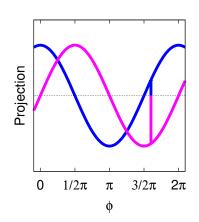
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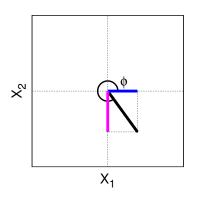


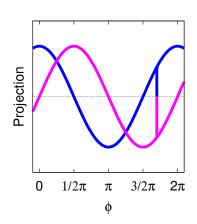
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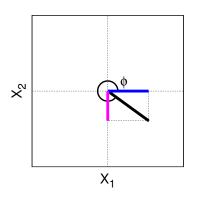


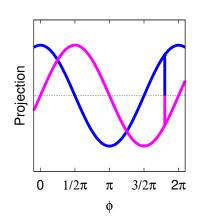
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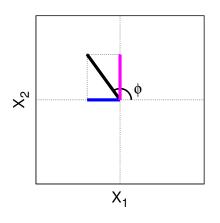
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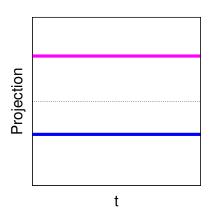




## Classical quadratures vs time in a rotating frame

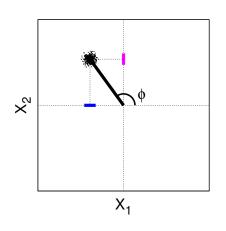
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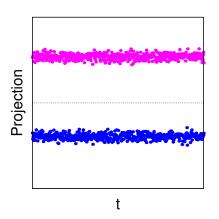




## Reality check quadratures vs time

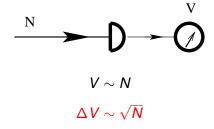
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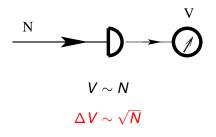
## Detector quantum noise

#### Simple photodetector

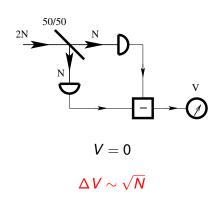


## Detector quantum noise

#### Simple photodetector



#### Balanced photodetector



## Transition from classical to quantum field

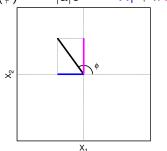
#### Classical analog

- Field amplitude a
- Field real part

$$X_1=(a^*+a)/2$$

• Field imaginary part  $X_2 = i(a^* - a)/2$ 

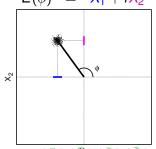
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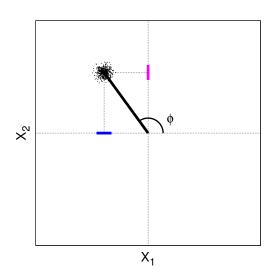
#### Quantum approach

- Field operator â
- Amplitude quadrature  $\hat{X}_1 = (\hat{a}^{\dagger} + \hat{a})/2$
- Phase quadrature  $\hat{X}_2 = i(\hat{a}^{\dagger} \hat{a})/2$

$$\hat{E}(\phi) = \hat{X_1} + i\hat{X_2}$$



## Quantum optics summary



Light consist of photons

• 
$$\hat{N} = a^{\dagger}a$$

Commutator relationship

• 
$$[a, a^{\dagger}] = 1$$

• 
$$[X_1, X_2] = i/2$$

**Detectors** measure

- number of photons  $\hat{N}$
- Quadratures  $\hat{X_1}$  and  $\hat{X_2}$

Uncertainty relationship

$$\Delta X_1 \Delta X_2 \ge 1/4$$

# Heisenberg uncertainty principle and its optics equivalent



## Heisenberg uncertainty principle

 $\Delta p \Delta x \geq \hbar/2$ 

The more precisely the POSITION is determined, the less precisely the MOMENTUM is known, and vice versa

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#### Optics equivalent

 $\Delta \phi \Delta N > 1$ 

The more precisely the PHASE is determined, the less precisely the AMPLITUDE is known, and vice versa

# Heisenberg uncertainty principle and its optics equivalent



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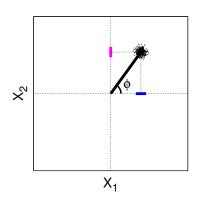
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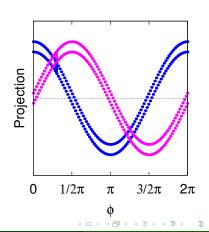
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#### Optics equivalent strict definition

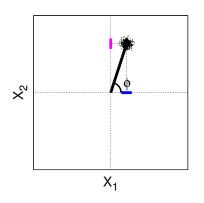
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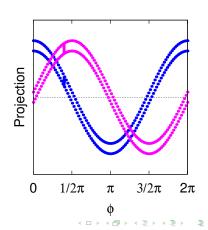
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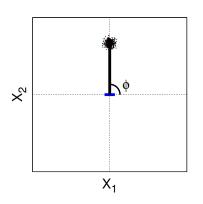


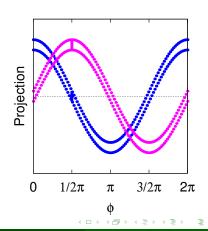
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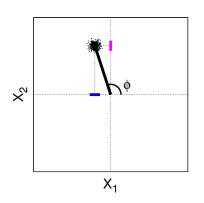


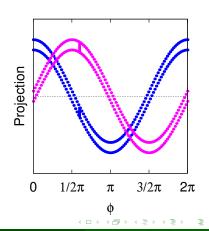
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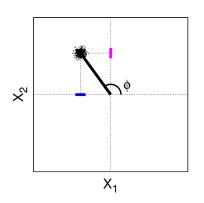


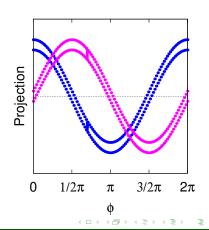
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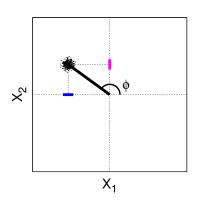


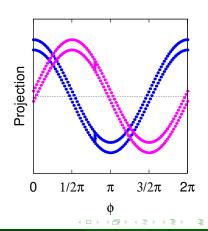
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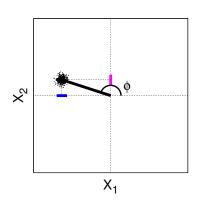


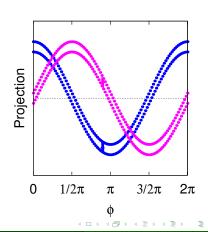
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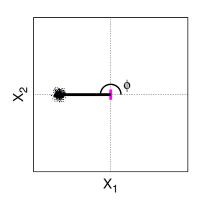


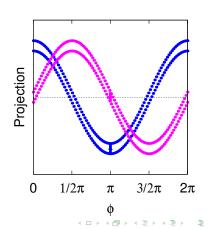
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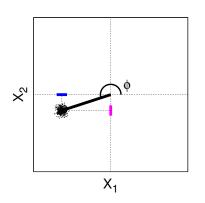


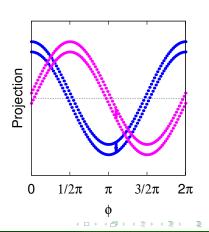
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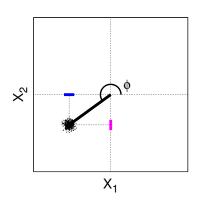


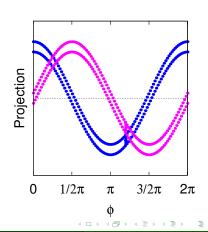
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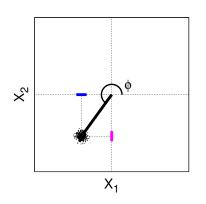


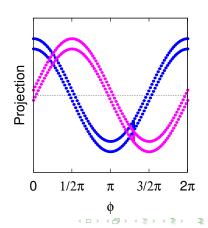
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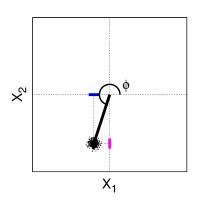


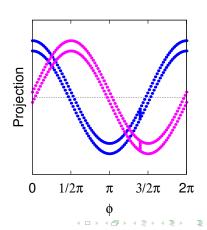
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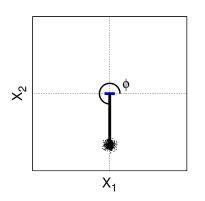


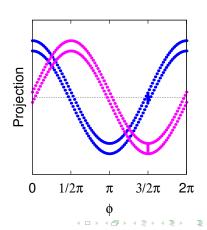
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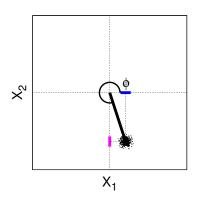


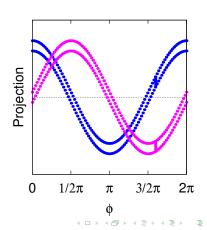
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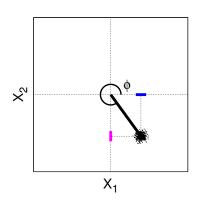


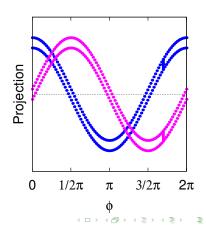
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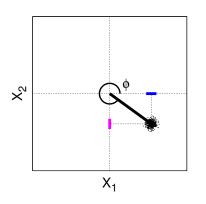


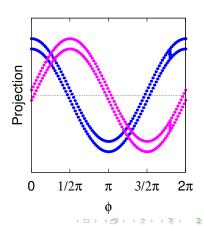
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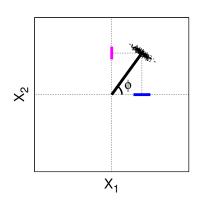


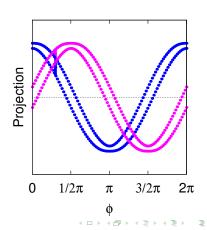
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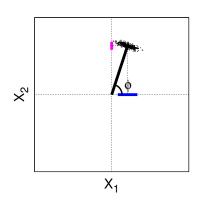


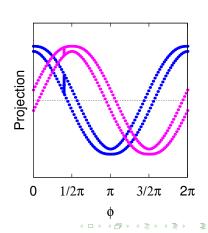
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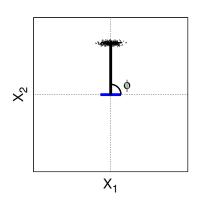


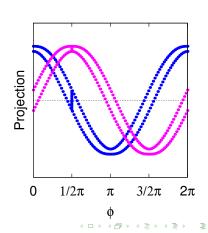
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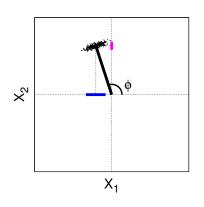


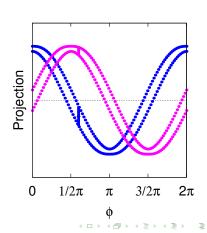
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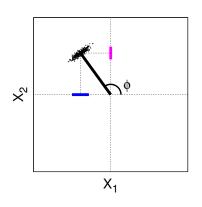


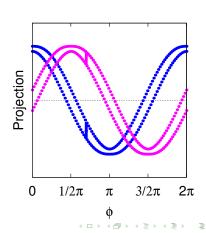
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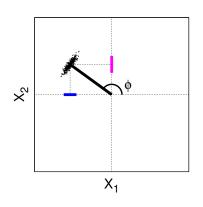


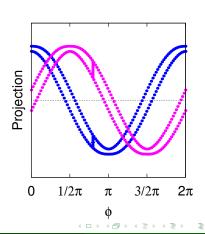
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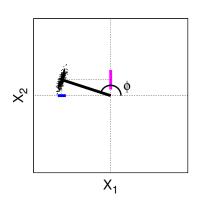


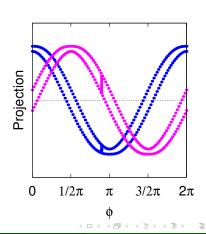
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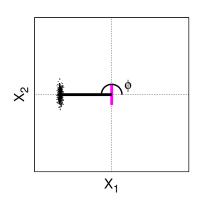


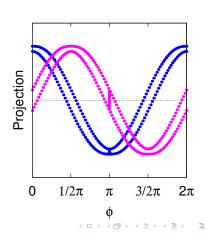
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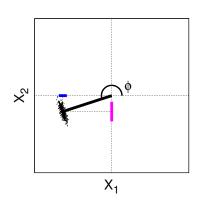


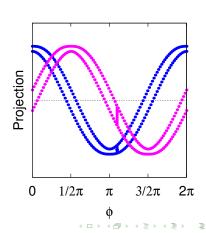
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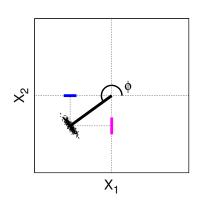


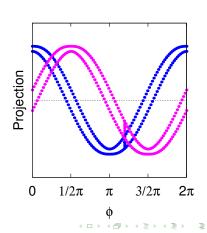
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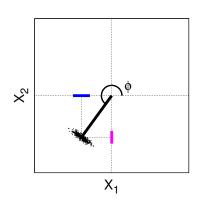


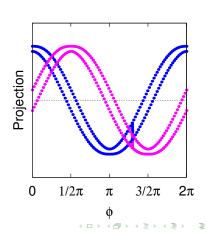
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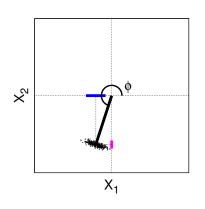


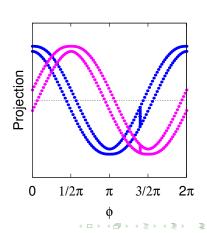
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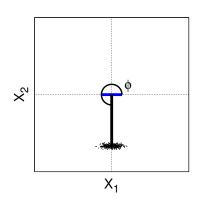


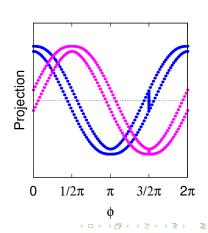
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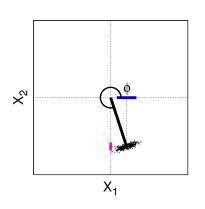


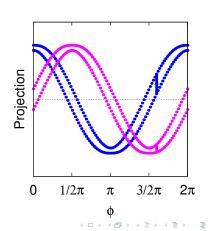
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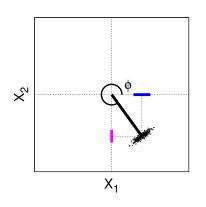


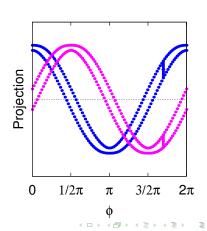
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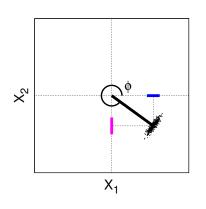


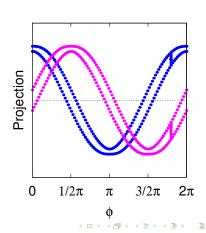
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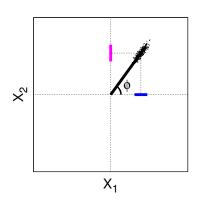


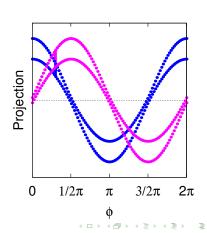
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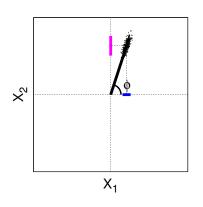


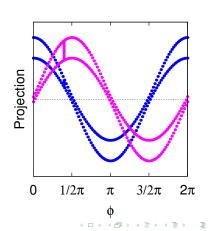
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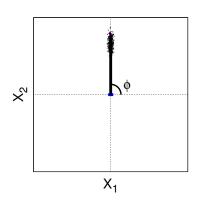


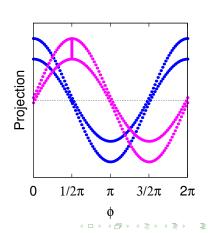
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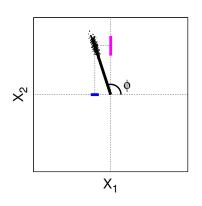


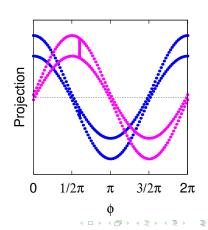
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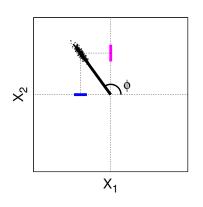


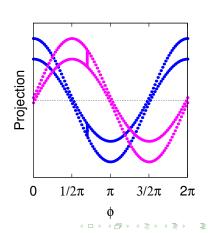
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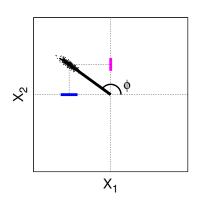


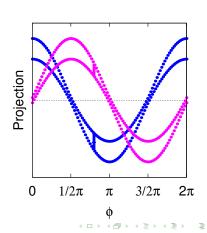
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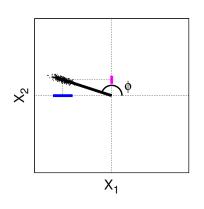


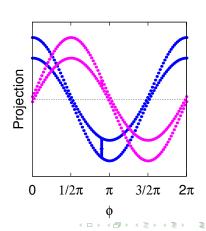
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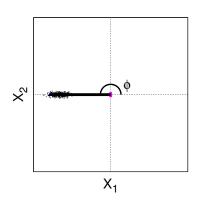


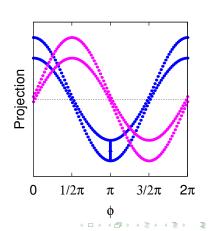
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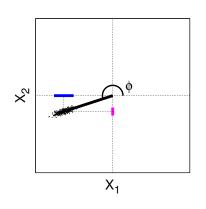


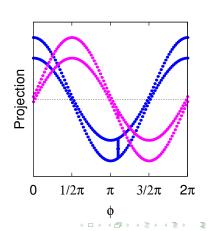
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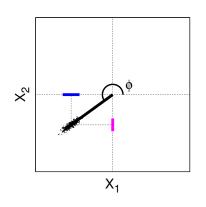


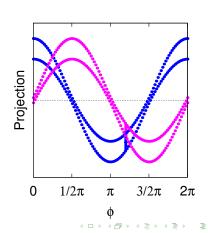
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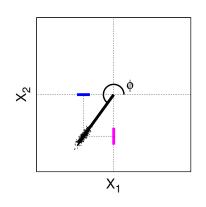


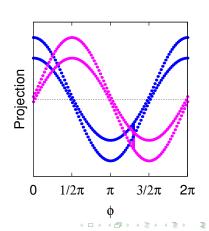
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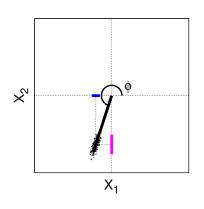


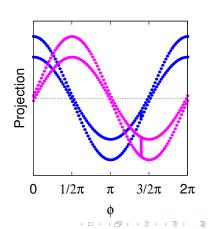
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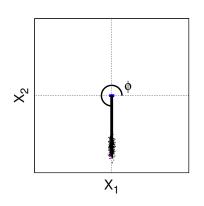


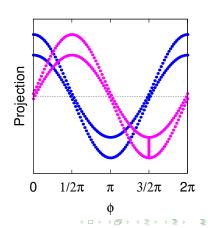
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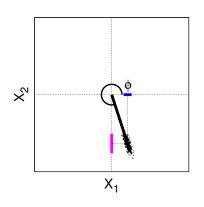


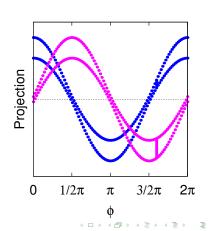
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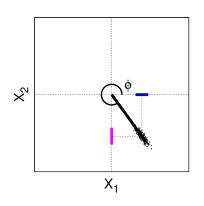


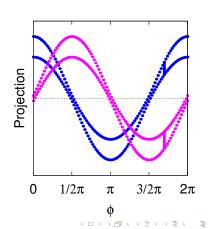
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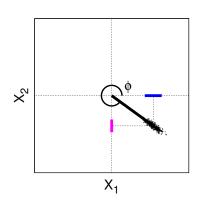


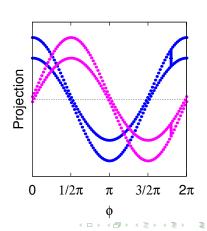
$$\Delta X_1 \Delta X_2 = 1/4$$

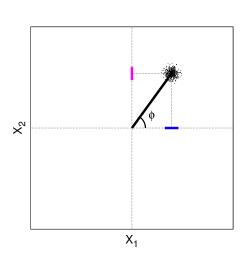




$$\Delta X_1 \Delta X_2 = 1/4$$

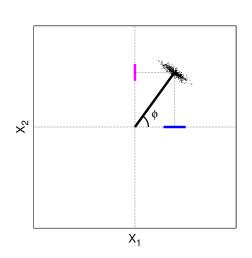




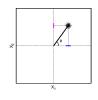


# Unsqueezed coherent



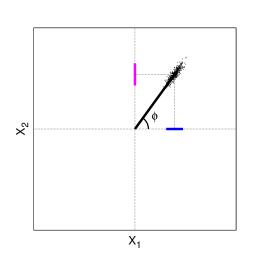


# Unsqueezed coherent



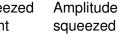
# Amplitude squeezed

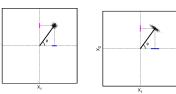




# coherent

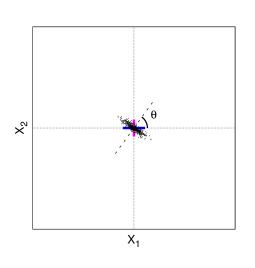






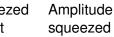
Phase squeezed





#### Unsqueezed coherent









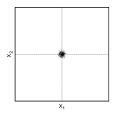
Phase squeezed

Vacuum squeezed



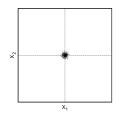


# Take a vacuum state |0>



$$H=\frac{1}{2}$$

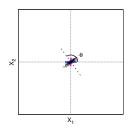
# Take a vacuum state |0>



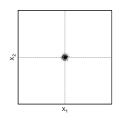
$$H=\frac{1}{2}$$

# Apply squeezing operator $|\xi>=\hat{S}(\xi)|0>$

$$\hat{S}(\xi) = e^{\frac{1}{2}\xi^*a^2 - \frac{1}{2}\xi a^{\dagger 2}}$$



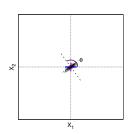
#### Take a vacuum state |0>



$$H=\frac{1}{2}$$

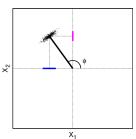
# Apply squeezing

$$\hat{S}(\xi) = e^{\frac{1}{2}\xi^*a^2 - \frac{1}{2}\xi a^{\dagger 2}}$$



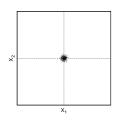
#### Apply displacement operator $|\xi\rangle = \hat{S}(\xi)|0\rangle$ operator $|\alpha, \xi\rangle = \hat{D}(\alpha)|s\rangle$

$$\hat{D}(\alpha) = e^{\alpha a^{\dagger} - \alpha^* a}$$



$$<\alpha, \xi | X_1 | \alpha, \xi > = Re(\alpha),$$
  
 $<\alpha, \xi | X_2 | \alpha, \xi > = Im(\alpha)$ 

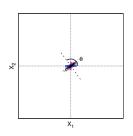
#### Take a vacuum state |0>



$$H=\frac{1}{2}$$

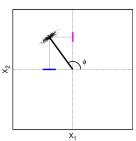
# Apply squeezing

$$\hat{S}(\xi) = e^{rac{1}{2}\xi^*a^2 - rac{1}{2}\xi a^{\dagger 2}}$$



#### Apply displacement operator $|\xi>=\hat{S}(\xi)|0>$ operator $|\alpha,\xi>=\hat{D}(\alpha)|s>$

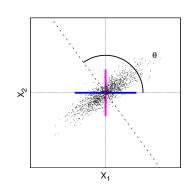
$$\hat{D}(\alpha) = e^{\alpha a^{\dagger} - \alpha^* a}$$



$$<\alpha, \xi | X_1 | \alpha, \xi > = Re(\alpha),$$
  
 $<\alpha, \xi | X_2 | \alpha, \xi > = Im(\alpha)$ 

Notice  $\Delta X_1 \Delta X_2 = \frac{1}{4}$ 

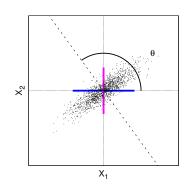
# Squeezed state $|\xi>=\hat{S}(\xi)|0>$ properties



$$\hat{S}(\xi) = e^{rac{1}{2}\xi^*a^2 - rac{1}{2}\xi a^{\dagger 2}}, \xi = re^{i heta}$$
 If  $heta = 0$  
$$< \xi |(\Delta X_1)^2|\xi> = rac{1}{4}e^{-2r}$$
 
$$< \xi |(\Delta X_2)^2|\xi> = rac{1}{4}e^{2r}$$

$$<\xi|(\Delta X_1)^2|\xi> = \frac{1}{4}(\cosh^2 r + \sinh^2 r - 2\sinh r\cosh r\cos\theta)$$
  
$$<\xi|(\Delta X_2)^2|\xi> = \frac{1}{4}(\cosh^2 r + \sinh^2 r + 2\sinh r\cosh r\cos\theta)$$

## Photon number of squeezed state $|\xi>$



Probability to detect given number of photons  $C = \langle n | \xi \rangle$  for squeezed vacuum

even

$$C_{2m} = (-1)\frac{\sqrt{(2m)!}}{2^m m!} \frac{(e^{i\theta} \tanh r)^m}{\sqrt{\cosh r}}$$

odd

$$C_{2m+1}=0$$

Average number of photons in general squeezed state

$$<\alpha,\xi|\mathbf{a}^{\dagger}\mathbf{a}|\alpha,\xi>=\alpha+\sinh^2r$$

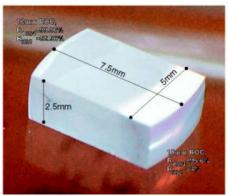
### Tools for squeezing

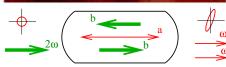
# Tools for squeezing



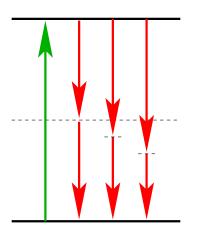
## Tools for squeezing







### Two photon squeezing picture

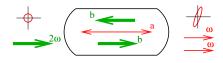


Squeezing operator

$$\hat{S}(\xi) = e^{\frac{1}{2}\xi^*a^2 - \frac{1}{2}\xi a^{\dagger 2}}$$

Parametric down-conversion in crystal

$$\hat{H} = i\hbar\chi^{(2)}(\mathbf{a}^2b^{\dagger} - \mathbf{a}^{\dagger2}b)$$



#### Squeezing

result of correlation of upper and lower sidebands

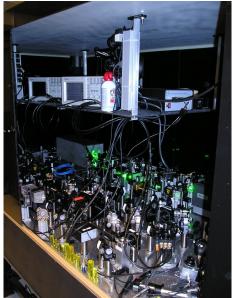
# Squeezer appearance

# Squeezer appearance

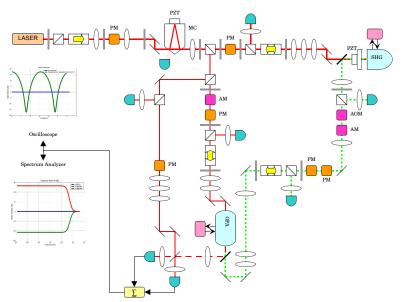


# Squeezer appearance





### Crystal squeezing setup scheme



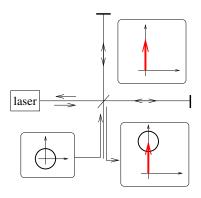
## Possible squeezing applications

- improvements any shot noise limited optical sensors
- noiseless signal amplification
- photon pair generation, entanglement, true single photon sources
- interferometers sensitivity boost (for example gravitational wave antennas)
- light free measurements
- quantum memory probe and information carrier

### Squeezing and interferometer

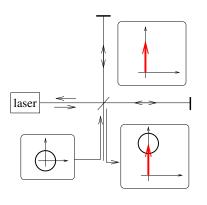
## Squeezing and interferometer

#### Vacuum input

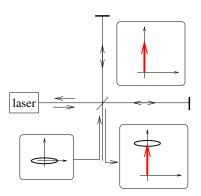


## Squeezing and interferometer

#### Vacuum input



#### Squeezed input

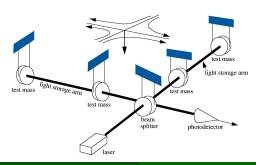


## Laser Interferometer Gravitational-wave Observatory

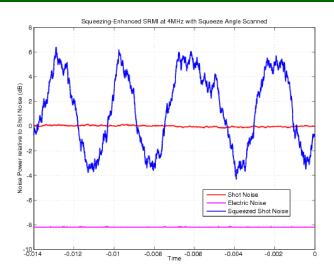




- L = 4 km
- $h \sim 2 \times 10^{-23}$
- $\bullet$   $\Delta L \sim 10^{-20} \text{ m}$

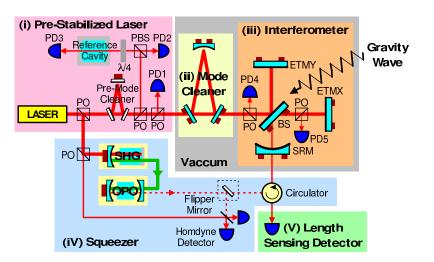


### Squeezing level vs time (unlocked)

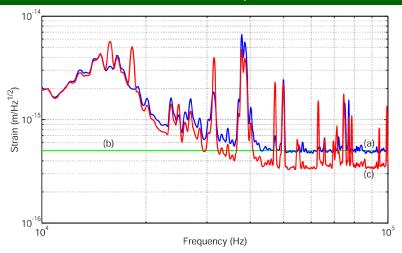


"A quantum-enhanced prototype gravitational-wave detector", Nature Physics, **4**, 472-476, (2008).

#### GW 40m detector and squeezer



## GW 40m detector with 4dB of squeezed vacuum



Signal to noise improvement by factor of 1.43

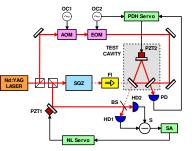


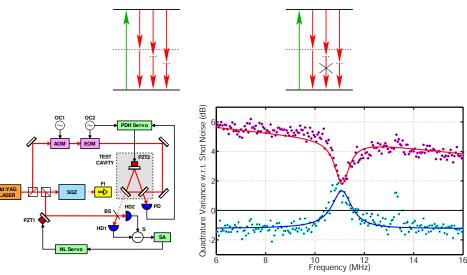












"Noninvasive measurements of cavity parameters by use of squeezed vacuum", Physical Review A, 74, 033817, (2006).

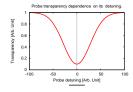
## Summary for crystal squeezing

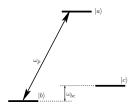
#### Pros

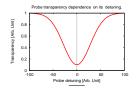
- mainstream: many different nonlinear crystals available
- so far the best squeezers
  - maximum squeezing value detected 11.5 dB at 1064 nm
  - Moritz Mehmet, Henning Vahlbruch, Nico Lastzka, Karsten Danzmann, and Roman Schnabel, "Observation of squeezed states with strong photon-number oscillations", Phys. Rev. A 81, 013814 (2010)
- well understood

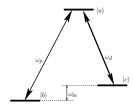
#### Cons

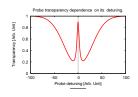
- crystals have limited transparency window
- thus squeezing is hard to generate at visible wavelength
  - at 795 nm only 4-6 dB squeezing is reported
- this limits applications of such squeezers for spectroscopy

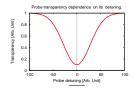


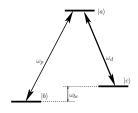


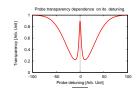




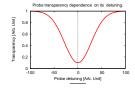


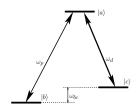


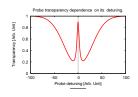




Storage and retrieval

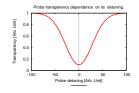


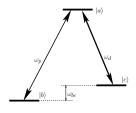


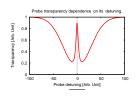


#### Storage and retrieval

single photon

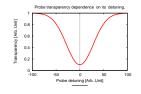


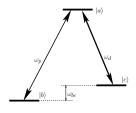


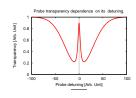


#### Storage and retrieval

- single photon
- squeezed state (Furusawa and Lvovsky PRL 100 2008)



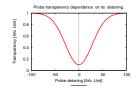


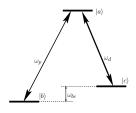


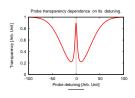
#### Storage and retrieval

- single photon
- squeezed state (Furusawa and Lvovsky PRL 100 2008)

Squeezed state requirements for a quantum memory probe





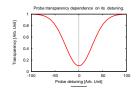


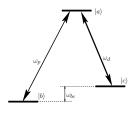
#### Storage and retrieval

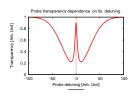
- single photon
- squeezed state (Furusawa and Lvovsky PRL 100 2008)

Squeezed state requirements for a quantum memory probe

- squeezing carrier at atomic wavelength (780nm, 795nm)
- squeezing within narrow resonance window at frequencies(<100kHz)</li>







#### Storage and retrieval

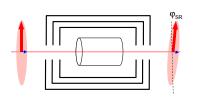
- single photon
- squeezed state (Furusawa and Lvovsky PRL 100 2008)

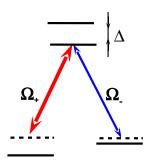
Squeezed state requirements for a quantum memory probe

- squeezing carrier at atomic wavelength (780nm, 795nm)
- squeezing within narrow resonance window at frequencies(<100kHz)</li>

Traditional nonlinear crystal based squeezers are capable of it, but they are extremely technically challenging especially at short wave length.

## Self-rotation of elliptical polarization in atomic medium

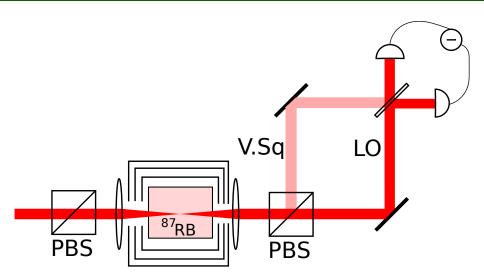




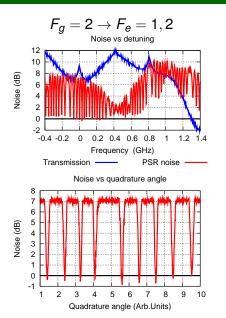
A.B. Matsko et al., PRA 66, 043815 (2002): theoretically prediction of 4-6 dB noise suppression

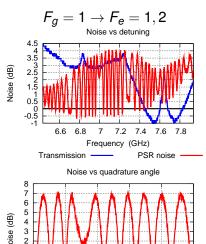
$$a_{out} = a_{in} + \frac{igL}{2}(a_{in}^{\dagger} - a_{in}) \tag{2}$$

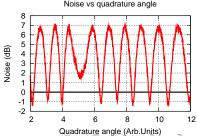
## Setup



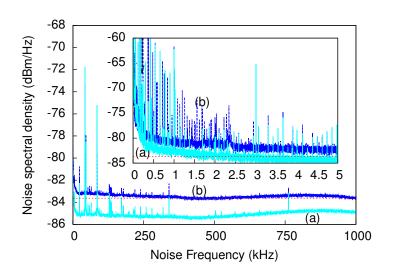
## Noise contrast vs detuning in hot 87Rb vacuum cell



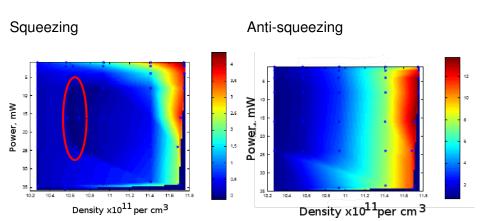




### Atomic low frequency squeezing source

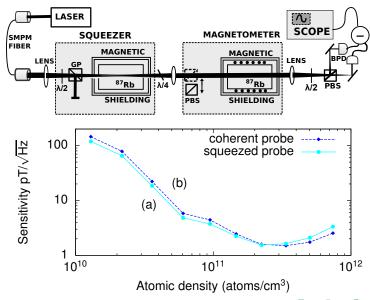


## Squeezing region



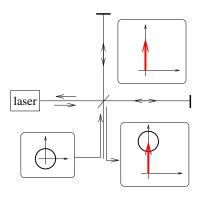
Observation of reduction of quantum noise below the shot noise limit is corrupted by the excess noise due to atomic interaction with atoms.

#### Magnetometer with squeezing enhancement

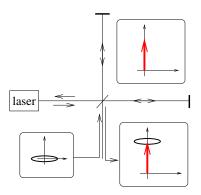


#### Quantum limited interferometers revisited

#### Vacuum input

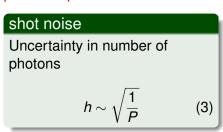


#### Squeezed input



Next generation of LIGO will be quantum optical noise limited at almost all detection frequencies.

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quantum optical noise limited at almost all detection frequencies.

#### shot noise

Uncertainty in number of photons

$$h \sim \sqrt{\frac{1}{P}}$$
 (3)

#### radiation pressure noise

Photons impart momentum to mirrors

$$h \sim \sqrt{\frac{P}{M^2 f^4}} \tag{4}$$

Next generation of LIGO will be

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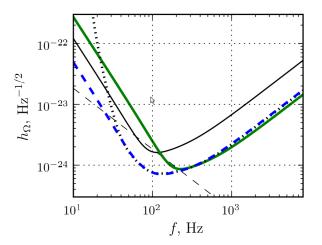
Photons impart momentum to mirrors

$$h \sim \sqrt{\frac{P}{M^2 f^4}} \tag{4}$$

There is no optimal light power to suit all detection frequency. Optimal power depends on desired detection frequency.

## Interferometer sensitivity improvement with squeezing

Projected advanced LIGO sensitivity

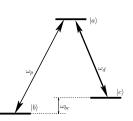


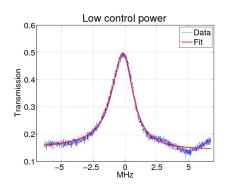
F. Ya. Khalili Phys. Rev. D 81, 122002 (2010)

$$\left(\begin{array}{c} V_1^{out} \\ V_2^{out} \end{array}\right) = \\ \left(\begin{array}{c} A_+^2 & A_-^2 \\ A_-^2 & A_+^2 \end{array}\right) \left(\begin{array}{c} V_1^{in} \\ V_2^{in} \end{array}\right) + \left[1 - \left(A_+^2 + A_-^2\right)\right] \left(\begin{array}{c} 1 \\ 1 \end{array}\right)$$

$$\varphi_{\pm} = \frac{1}{2} (\Theta_{+} \pm \Theta_{-})$$

$$A_{\pm} = \frac{1}{2} (T_{+} \pm T_{-})$$

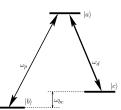


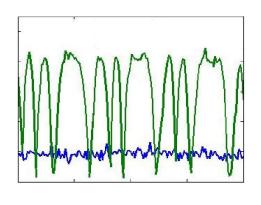


$$\left(\begin{array}{c} V_1^{out} \\ V_2^{out} \end{array}\right) = \\ \left(\begin{array}{c} A_+^2 & A_-^2 \\ A_-^2 & A_+^2 \end{array}\right) \left(\begin{array}{c} V_1^{in} \\ V_2^{in} \end{array}\right) + \left[1 - \left(A_+^2 + A_-^2\right)\right] \left(\begin{array}{c} 1 \\ 1 \end{array}\right)$$

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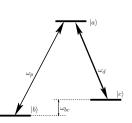


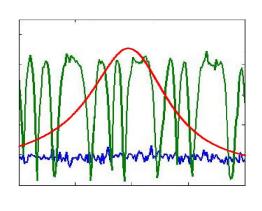


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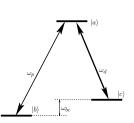


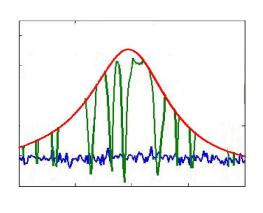


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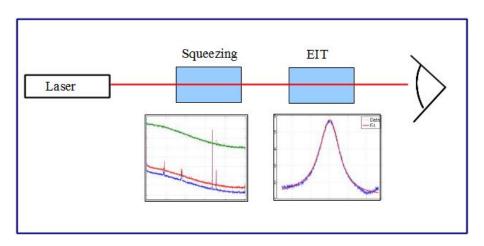
$$\varphi_{\pm} = \frac{1}{2} (\Theta_{+} \pm \Theta_{-})$$

$$A_{\pm} = \frac{1}{2} (T_{+} \pm T_{-})$$

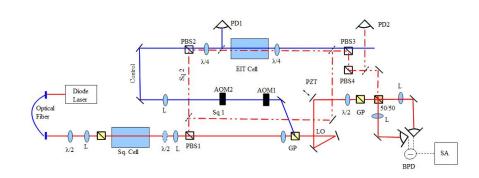




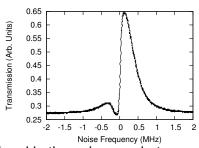
#### Squeezing and EIT filter setup



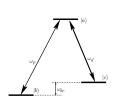
#### Squeezing and EIT filter setup



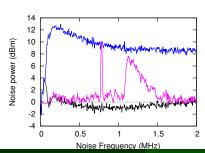
### EIT filter and measurements without light

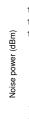


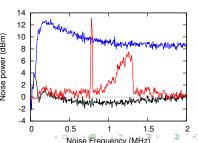
#### Coherent signal



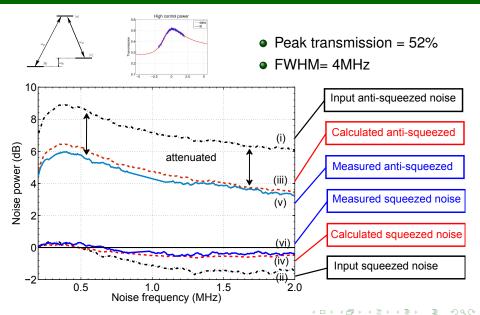
Signal in the noise quadratures



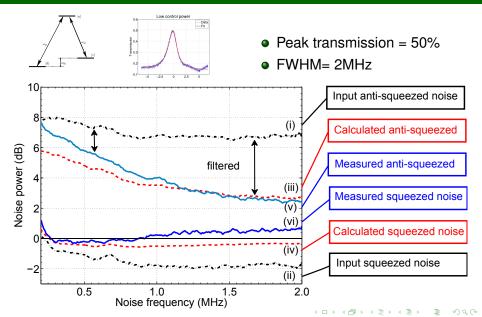




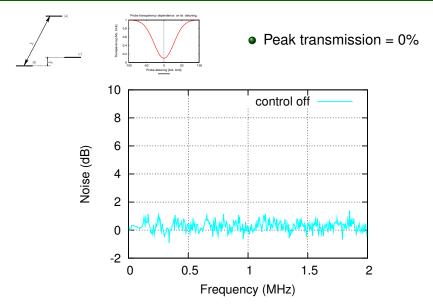
#### Wide EIT filter and squeezing



### Narrow EIT filter and squeezing

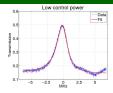


#### Control off no EIT and no squeezing at the output



### Squeezing angle rotation

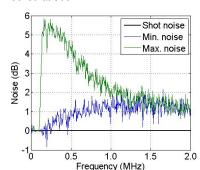




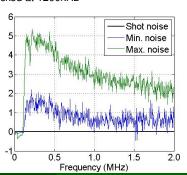


$$\left( \begin{array}{c} V_1^{out} \\ V_2^{out} \end{array} \right) = \\ \left( \begin{array}{c} \cos^2\varphi_+ & \sin^2\varphi_+ \\ \sin^2\varphi_+ & \cos^2\varphi_+ \end{array} \right) \left( \begin{array}{cc} A_+^2 & A_-^2 \\ A_-^2 & A_+^2 \end{array} \right) \left( \begin{array}{c} V_1^{in} \\ V_2^{in} \end{array} \right) + \left[ 1 - \left( A_+^2 + A_-^2 \right) \right] \left( \begin{array}{c} 1 \\ 1 \end{array} \right)$$

#### Locked at 300kHz

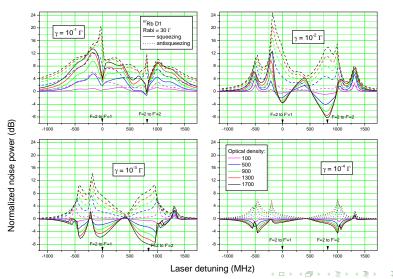


#### Locked at 1200kHz



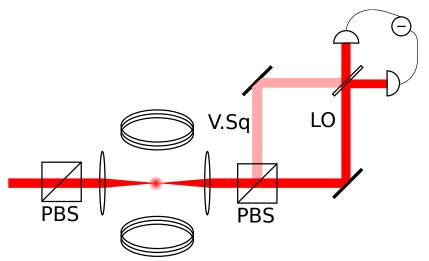
## Theoretical prediction for MOT squeezing with 87Rb

 $F_g = 2 \rightarrow F_e = 1,2$  high optical density is very important

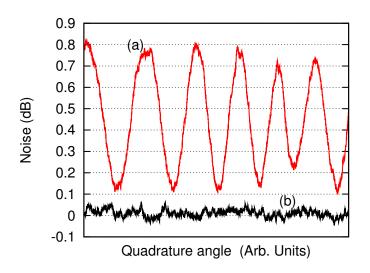


#### MOT squeezer

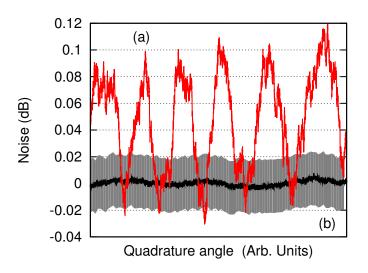
Cloud size =1 mm, T = 200  $\mu$ K, N = 7  $\times$  10<sup>9</sup> 1/cm<sup>3</sup>, OD = 2, beam size = 0.1 mm, 10<sup>5</sup> interacting atoms



#### Noise contrast in MOT with $^{87}$ Rb $F_q = 2 \rightarrow F_e = 1$

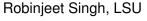


# Squeezing in MOT with $^{87}$ Rb $F_g=2 ightarrow F_e=1$



## People

Travis Horrom and Gleb Romanov





Irina Novikova

Jonathan P. Dowling, LSU





#### Summary

- Squeezing is exciting
- many applications benefit from squeezing
- there is still a lot of interesting physics to do

Support from

