

It will be useful to define some notation. We will use a bar on a 2-index tensor to mean the following combination:

$$\bar{A}_{\mu\nu} \equiv \frac{1}{2}(A_{\mu\nu} + A_{\nu\mu}) - \frac{1}{2}\gamma_{\mu\nu} A^{\sigma}_{\sigma}$$

For a symmetric tensor like  $h_{\mu\nu}$ ,

$$\bar{h}_{\mu\nu} = h_{\mu\nu} - \frac{1}{2}\gamma_{\mu\nu} h^{\sigma}_{\sigma}$$

Note:  $\bar{h}_{\mu\nu} = h_{\mu\nu}$

More notation:  $h \equiv h^{\sigma}_{\sigma}$  "Trace",  $\bar{h}^{\sigma}_{\sigma} = -h$

The equations of motion for  $h_{\alpha\beta}$  take the form

$$\begin{aligned} G_{\alpha\beta} &\equiv -\partial_{\sigma}\partial^{\sigma}h_{\alpha\beta} + (\partial_{\alpha}\partial^{\sigma}h_{\beta\sigma} + \partial_{\beta}\partial^{\sigma}h_{\alpha\sigma}) \\ &\quad - (\partial_{\alpha}\partial_{\beta}h + \gamma_{\alpha\beta}\partial_{\mu}\partial^{\mu}h^{mn}) + \gamma_{\alpha\beta}\partial_{\mu}\partial^{\mu}h \\ &= -\lambda T_{\alpha\beta} \end{aligned}$$

Taking the trace:

$$\begin{aligned} G &\equiv -\partial_{\sigma}\partial^{\sigma}h + 2\partial^{\alpha}\partial^{\sigma}h_{\alpha\sigma} - \partial_{\alpha}\partial^{\alpha}h + 4\partial_{\mu}\partial^{\mu}h - 4\partial_{\mu}\partial^{\mu}h^{mn} \\ &= 2(\partial^{\alpha}\partial^{\sigma}h_{\alpha\sigma} + \partial_{\alpha}\partial^{\alpha}h) \\ &= -\lambda T \end{aligned}$$

$$\begin{aligned} \bar{G}_{\alpha\beta} &= G_{\alpha\beta} - \frac{1}{2} \eta_{\alpha\beta} G \\ &= -2\partial_\alpha \partial^\sigma h_{\alpha\beta} + 2\partial_\alpha \partial^\sigma h_{\alpha\sigma} + 2\partial^\sigma \partial_\alpha h_{\beta\sigma} - \partial_\alpha \partial_\beta h \\ &= -2\bar{T}_{\alpha\beta} = -\lambda \left( T_{\alpha\beta} - \frac{1}{2} \eta_{\alpha\beta} T \right) \end{aligned}$$

These equations are invariant under the transformations  $h_{\mu\nu} \rightarrow h_{\mu\nu} + \partial_\mu \xi_\nu + \partial_\nu \xi_\mu$  for arbitrary functions  $\xi_\mu(x)$ :

$$\begin{aligned} \bar{G}_{\alpha\beta} &\rightarrow \bar{G}_{\alpha\beta} - 2\partial_\alpha \partial^\sigma (\partial_\beta \xi_\sigma + \partial_\sigma \xi_\beta) + 2\partial_\alpha \partial^\sigma (\partial_\sigma \xi_\alpha + \partial_\alpha \xi_\sigma) \\ &\quad + 2\partial^\sigma \partial_\alpha (\partial_\beta \xi_\sigma + \partial_\sigma \xi_\beta) - \partial_\alpha \partial_\beta (2\partial_\sigma \xi^\sigma) \\ &= \bar{G}_{\alpha\beta} \end{aligned}$$

Given a conserved  $T_{\alpha\beta}$ , for any solution to the equations for  $h_{\alpha\beta}$  there are infinitely more solutions given by this transformation.

This invariance is reminiscent of gauge invariance in the relativistic formulation of electrodynamics. The equations of motion of electrodynamics can be written in terms of the field strength  $F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$ .

The field strength is invariant under  $A_\mu \rightarrow A_\mu + \partial_\mu \Theta(x)$  for arbitrary  $\Theta(x)$ . Physical observables are independent of  $\Theta(x)$ , so the gauge invariance reflects a redundancy in the description of the field  $A_\mu(x)$ .

Just as in electrodynamics, we can choose to fix the gauge by a condition on  $h_{\mu\nu}$  that makes certain equations look simpler.

For example, we may choose to set

$$\boxed{\partial_{\mu} h^{\mu\nu} = \frac{1}{2} \partial^{\nu} h^{\mu\mu}} \quad \text{Harmonic gauge.}$$

To see that we can make this choice, assume

$$\partial_{\mu} h^{\mu\nu} - \frac{1}{2} \partial^{\nu} h = f^{\nu} \neq 0.$$

$$\text{Let } h^{\mu\nu} \rightarrow h^{\mu\nu} + \partial^{\mu} \xi^{\nu} + \partial^{\nu} \xi^{\mu}$$

$$\begin{aligned} \partial_{\mu} h^{\mu\nu} - \frac{1}{2} \partial^{\nu} h &\rightarrow \partial_{\mu} h^{\mu\nu} - \frac{1}{2} \partial^{\nu} h + \partial_{\mu} \partial^{\mu} \xi^{\nu} + \partial_{\mu} \partial^{\nu} \xi^{\mu} - \frac{1}{2} \cdot 2 \partial^{\nu} \partial_{\mu} \xi^{\mu} \\ &= 0 \quad \text{if } \partial_{\mu} \partial^{\mu} \xi^{\nu} = -f^{\nu}, \end{aligned}$$

which can be solved for  $\xi^{\nu}$ .

Note that the harmonic gauge condition does not completely fix the gauge, because we can always add a solution to  $\partial_{\mu} \partial^{\mu} \xi^{\nu} = 0$ , the wave equation.

In the harmonic gauge, the equations of motion simplify to

$$\boxed{-\partial_{\alpha} \partial^{\alpha} h_{\alpha\beta} = -\lambda \bar{T}_{\alpha\beta}}$$

of Weinberg  
Ch. 10

## Plane wave solutions

Consider the homogeneous equations

$$(*) \quad \begin{cases} \partial_\alpha \partial^\alpha h_{\mu\nu} = 0 \\ \partial_\mu h^\mu{}_\nu = \frac{1}{2} \partial_\nu h^\mu{}_\mu \end{cases}$$

Plane wave solutions have the form

$$h_{\mu\nu}(x) = \epsilon_{\mu\nu} \exp(i k \cdot x) + \epsilon_{\mu\nu}^* \exp(-i k \cdot x)$$

where  $k \cdot x \equiv k_\mu x^\mu$ ,  $\epsilon_{\mu\nu} \equiv$  polarization tensor  
(in general complex)

The equations (\*) imply

$$\begin{cases} -k_\alpha k^\alpha h_{\mu\nu} = 0 \\ k_\mu \epsilon^\mu{}_\nu = \frac{1}{2} k_\nu \epsilon^\mu{}_\mu \end{cases}$$

Symmetry of  $h_{\mu\nu}$  implies  $\epsilon_{\mu\nu} = \epsilon_{\nu\mu}$

Consider the gauge transformations  $h_{\mu\nu} \rightarrow h_{\mu\nu} + \partial_\mu \xi_\nu + \partial_\nu \xi_\mu$   
with  $\xi^\mu(x) = i \epsilon^\mu \exp(i k \cdot x) + i \epsilon^{\mu*} \exp(-i k \cdot x)$

This is equivalent to transforming the polarizations

$$\epsilon_{\mu\nu} \rightarrow \epsilon_{\mu\nu} + k_\mu \xi_\nu + k_\nu \xi_\mu$$

This class of gauge transformations preserves the harmonic conditions.

$$\begin{aligned} k_\mu \epsilon^\mu{}_\nu &\rightarrow k_\mu \epsilon^\mu{}_\nu + \overset{0}{k} \cdot \vec{k} \xi_\nu + (k \cdot \epsilon) k_\nu \\ \frac{1}{2} k_\nu \epsilon^\mu{}_\mu &\rightarrow \frac{1}{2} k_\nu \epsilon^\mu{}_\mu + (k \cdot \epsilon) k_\nu \end{aligned}$$

Number of independent propagating solutions:

For each  $k^\mu$  satisfying  $k_\mu k^\mu = 0$ ,

$E_{\mu\nu}$  — symmetric  $4 \times 4$  matrix  $\rightarrow 10$  components

harmonic conditions  $-4$

remaining gauge freedom  $-4$

2 independent polarizations

Example: wave traveling in  $x^3$ -direction.

$$k^1 = k^2 = 0, \quad k^3 = k^0 = k > 0.$$

$$\left. \begin{aligned} \text{Harmonic conditions: } k^3 \epsilon_{31} + k^0 \epsilon_{01} &= k^3 \epsilon_{32} + k^0 \epsilon_{02} = 0 \\ k^3 \epsilon_{33} + k^0 \epsilon_{03} &= (k^3 \epsilon_{30} + k^0 \epsilon_{00}) \\ &= \frac{1}{2} k (\epsilon_{11} + \epsilon_{22} + \epsilon_{33} - \epsilon_{00}) \end{aligned} \right\}$$

$$k^3 = k^0 = k$$

$$\Rightarrow \epsilon_{31} + \epsilon_{01} = \epsilon_{32} + \epsilon_{02} = 0$$

$$\epsilon_{33} + \epsilon_{03} = -(\epsilon_{30} + \epsilon_{00}) = \frac{1}{2} (\epsilon_{11} + \epsilon_{22} + \epsilon_{33} - \epsilon_{00})$$

$$\Rightarrow \left. \begin{aligned} \epsilon_{01} &= -\epsilon_{31}, & \epsilon_{02} &= -\epsilon_{32}, \\ \epsilon_{22} &= -\epsilon_{11}, & \epsilon_{03} &= -\frac{1}{2} (\epsilon_{33} + \epsilon_{00}) \end{aligned} \right\}$$

$\epsilon_{01}, \epsilon_{02}, \epsilon_{22}, \epsilon_{03}$   
dependent on other polarizations

Residual gauge freedom:  $\epsilon_{13} \rightarrow \epsilon_{13} + k \xi_1$

$$\epsilon_{23} \rightarrow \epsilon_{23} + k \xi_2$$

$$\epsilon_{33} \rightarrow \epsilon_{33} + 2k \xi_3$$

$$\epsilon_{00} \rightarrow \epsilon_{00} - 2k \xi_0$$

absolute  
Do not have physical significance.

$\Rightarrow$  Only two components ( $\epsilon_{11}, \epsilon_{12}$ ) have absolute physical significance.

## Helicity of Gravitational Waves

Consider a rotation by angle  $\theta$  about the  $x^3$ -axis,

$$\Lambda^{\mu}_{\nu} = \begin{pmatrix} 1 & & & \\ & \cos\theta & -\sin\theta & \\ & +\sin\theta & \cos\theta & \\ & & & 1 \end{pmatrix}, \quad (\Lambda^{-1})^{\mu}_{\nu} = \begin{pmatrix} 1 & & & \\ & \cos\theta & \sin\theta & \\ & -\sin\theta & \cos\theta & \\ & & & 1 \end{pmatrix}$$

The polarization transforms as  $E_{\mu\nu} \rightarrow (\Lambda^{-1})^{\alpha}_{\mu} (\Lambda^{-1})^{\beta}_{\nu} E_{\alpha\beta} \equiv E'_{\mu\nu}$ .

$$\text{Defining } e_{\pm} \equiv e_{11} \mp i e_{12} = -e_{22} \mp i e_{21}$$

$$f_{\pm} \equiv e_{31} \mp i e_{32} = -e_{01} \mp i e_{02}$$

It is straight forward to check that under the rotation,

$$e'_{\pm} = \exp(\pm 2i\theta) e_{\pm} \quad \leftarrow \text{helicity } \pm 2$$

$$f'_{\pm} = \exp(\pm i\theta) f_{\pm} \quad \leftarrow \text{helicity } \pm 1$$

$$e'_{33} = e_{33}, \quad e'_{00} = e_{00} \quad \leftarrow \text{helicity } 0$$

Any plane wave which transforms as  $\psi' = e^{ih\theta} \psi$  under a rotation by  $\theta$  about the direction of motion is said to have helicity  $h$ .

Gravitational waves are decomposed into helicity 2, 1, 0 parts, but only the helicity 2 parts are physically significant as per our previous discussion.

Comparison w/ electromagnetic waves: In Lorenz gauge  $\begin{cases} \partial_{\mu} \partial^{\mu} A_{\alpha} = 0 \\ \partial_{\alpha} A^{\alpha} = 0 \end{cases}$

Residual gauge freedom:  $A^{\mu} \rightarrow A^{\mu} + \partial^{\mu} \theta$  s.t.  $\partial_{\mu} \partial^{\mu} \theta = 0$

# propagating degrees of freedom =  $4 - 1 - 1 = 2$

← Lorenz gauge condition      ← residual gauge freedom

physical plane waves have helicity  $\pm 1$  in electromagnetism.