## Physics 786, Fall 2018 Problem Set 8, Due Wednesday, November 7.

## 1. Newtonian Stars

Consider static, spherically symmetric solutions to Einstein's equations for a fluid with density  $\rho(r)$  and pressure p(r), with metric of the form,

$$ds^{2} = -e^{2\phi(r)}dt^{2} + e^{2\lambda(r)}dr^{2} + r^{2}\left(d\theta^{2} + \sin^{2}\theta \,d\varphi^{2}\right).$$

Assume the fluid is nonrelativistic, and consider the nonrelativistic limit of Einstein's equations for this system. Assume  $\phi(0) = \lambda(0) = 0$ , and

$$\rho(r) = \rho_0 \left( 1 - \frac{r^2}{R^2} \right)$$

for  $r \leq R$ , and  $\rho(r) = 0$  for r > R.

a) Find the spacetime metric for r < R and r > R. The metric should be continuous across r = R.

b) Find the pressure p(r) in the star such that p(R) = 0.

## 2. The Planck Mass and Planck Length

For a particle of mass M, quantum mechanics is important at distance scales of order the Compton wavelength of the particle. What is the value of M such that the reduced Compton wavelength  $\lambda_C = \hbar/(Mc)$  is equal to the Schwarzschild radius of the particle? What is the corresponding value of  $\lambda_C$ ? Express your results in SI units.

The Planck mass is defined by  $M_{\rm Pl} = \sqrt{\hbar c/G}$ . Compare the value of M that you found with the Planck mass.