

## Review of Lagrangian Mechanics

Suppose a system is described by a set of  $N$  coordinates  $q_i(t)$ ,  $i=1, \dots, N$ .

Action Principle: There exists a functional of  $q_i$ ,  $\dot{q}_i$  and  $t$ , called the action, which is stationary about variations  $\delta q_i(t)$  along a classical path.

Lagrangian  $L(q_i, \dot{q}_i, t)$

Action  $S = \int_{t_1}^{t_2} dt L(q_i, \dot{q}_i, t)$

Variation  $q_i(t) \rightarrow q_i(t) + \delta q_i(t)$

$\delta q_i(t_1) = \delta q_i(t_2) = 0$  (Boundaries fixed)

$$\delta S = \int_{t_1}^{t_2} dt \sum_i \left[ \frac{\partial L}{\partial q_i} \delta q_i + \frac{\partial L}{\partial \dot{q}_i} \delta \dot{q}_i \right]$$

$$= \int_{t_1}^{t_2} dt \sum_i \delta q_i \left[ \frac{\partial L}{\partial q_i} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_i} \right]$$

$\delta S = 0$  for arbitrary small variations  $\delta q_i(t)$  along a classical path

$$\Rightarrow \boxed{\frac{\partial L}{\partial q_i} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_i} = 0} \quad \text{Euler-Lagrange Eqs.}$$

## Hamiltonian Formulation

Given a Lagrangian  $L$ , define the canonical momenta

$$p_i \equiv \frac{\partial L}{\partial \dot{q}_i} \quad \rightsquigarrow \quad \frac{dp_i}{dt} = \frac{\partial L}{\partial q_i} \quad \text{from Euler-Lagrange Eqs.}$$

Define the Hamiltonian  $H(p_i, q_i, t) \equiv \sum_i p_i \dot{q}_i - L$

$H$  must be written in terms of  $p$ 's and  $q$ 's, not the  $\dot{q}$ 's, and the  $p$ 's and  $q$ 's must be independent. (This is not always possible!)

Vary the coordinates and momenta:

$$\begin{aligned} dH &= \sum_i \left( dp_i \dot{q}_i + p_i \cancel{d\dot{q}_i} - \frac{\partial L}{\partial q_i} dq_i - \frac{\partial L}{\partial \dot{q}_i} \cancel{d\dot{q}_i} \right) \\ &= \sum_i \left( \dot{q}_i dp_i - \dot{p}_i dq_i \right) \quad \text{using the Euler-Lagrange Eqs.} \\ &= \sum_i \left( \frac{\partial H}{\partial p_i} dp_i + \frac{\partial H}{\partial q_i} dq_i \right) \end{aligned}$$

$$\rightsquigarrow \quad \boxed{\frac{\partial H}{\partial p_i} = \dot{q}_i, \quad \frac{\partial H}{\partial q_i} = -\dot{p}_i} \quad \text{Hamilton's Eqs.}$$

$$\begin{aligned} \text{If } \frac{\partial L}{\partial t} = 0 \rightsquigarrow \frac{dH}{dt} &= \sum_i \left( \frac{\partial H}{\partial p_i} \dot{p}_i + \frac{\partial H}{\partial q_i} \dot{q}_i \right) - \cancel{\frac{\partial L}{\partial t}} \\ &= \sum_i \left( \dot{q}_i \dot{p}_i - \dot{p}_i \dot{q}_i \right) = 0 \end{aligned}$$

In that case  $H$  is called the energy, and is conserved.

## Symmetries and Conservation Laws

Noether's Theorem: For every global symmetry parametrized by a continuous parameter there is a corresponding conservation law.

Under a variation of  $q_i(t)$ ,

$$\begin{aligned} L &\rightarrow L + \sum_i \left( \frac{\partial L}{\partial q_i} \delta q_i + \frac{\partial L}{\partial \dot{q}_i} \delta \dot{q}_i \right) \\ &= L + \sum_i \left( \dot{p}_i \delta q_i + p_i \delta \dot{q}_i \right) \quad \text{using the E-L Eqs.} \\ &= L + \frac{d}{dt} \left( \sum_i p_i \delta q_i \right) \quad \text{using } \delta \dot{q}_i = \frac{d}{dt} \delta q_i \end{aligned}$$

Now consider a class of variations parametrized by a continuous parameter  $\lambda$ :  $\delta q_i(t, \lambda) = \lambda \left. \frac{\partial q_i}{\partial \lambda} \right|_{\lambda=0}$  as  $\lambda \rightarrow 0$ .

Suppose under this class of transformations

$$L \rightarrow L + \lambda \frac{dF}{dt} \quad \text{for some } F(q_i, \dot{q}_i, t), \quad F(t_2) - F(t_1) = 0. \\ \text{(not using the eqs. of motion)}$$

Then the class of transformations  $q_i(t) \rightarrow q_i(t, \lambda)$  is called a symmetry.

$$\begin{aligned} \text{Hamilton's principle: } 0 = \delta S &= \int_{t_1}^{t_2} dt \delta L = \int_{t_1}^{t_2} dt \lambda \frac{dF}{dt} \\ &= \lambda (F(t_2) - F(t_1)) = 0 \end{aligned}$$

Such a symmetry transformation does not change the action or the equations of motion.

$$\begin{aligned} \text{Around the eqs. of motion } L &\rightarrow L + \frac{d}{dt} \left( \sum_i p_i \delta q_i(t, \lambda) \right) \\ &= L + \lambda \frac{d}{dt} \left( \sum_i p_i \frac{\partial q_i}{\partial \lambda} \right) \Big|_{\lambda=0} \text{ for } \lambda \ll 1 \\ &= L + \lambda \frac{dF}{dt} \text{ by assumption.} \end{aligned}$$

$$\Rightarrow \boxed{\frac{d}{dt} \left( \sum_i p_i \frac{\partial q_i}{\partial \lambda} \Big|_{\lambda=0} - F(q_i, \dot{q}_i, t) \right) = 0}$$

The quantity  $Q \equiv \sum_i p_i \frac{\partial q_i}{\partial \lambda} \Big|_{\lambda=0} - F(q_i, \dot{q}_i, t)$

is conserved.

Example: Space translation of point particles

$$\text{Lagrangian } L = \sum_i \frac{1}{2} m_i \dot{\vec{r}}_i^2 - \sum_{i,j} V_{ij}(|\vec{r}_i - \vec{r}_j|)$$

$\vec{r}_i \rightarrow \vec{r}_i + \vec{a} \lambda$  takes  $L \rightarrow L$  for any fixed  $\vec{a}$ .

$$\frac{\partial F_i(\lambda)}{\partial \lambda} = \vec{a}, \quad F=0, \quad \vec{p}_i = \frac{\partial L}{\partial \dot{\vec{r}}_i} = m_i \dot{\vec{r}}_i$$

$Q = \sum_i \vec{p}_i \cdot \vec{a}$  is conserved. Since this is true  $\forall \vec{a}$ ,  
 $\sum_i \vec{p}_i$  is conserved.

Translation invariance  $\rightarrow$  Momentum Conservation.

Example: Time translations  $q_i(t) \rightarrow q_i(t+\lambda)$

$$\left. \frac{\partial \mathcal{L}(t, \lambda)}{\partial \lambda} \right|_{\lambda=0} = \frac{d\mathcal{L}(t)}{dt}$$

$$\delta L = \lambda \frac{dL}{dt} \rightsquigarrow F=L$$

$$Q = \sum_i p_i \dot{q}_i - L = H \quad \text{is conserved.}$$

Time translation invariance  $\rightarrow$  Energy conservation

## Classical Field Theory

$$\begin{aligned} q_i(t) &\rightsquigarrow \phi_i(\vec{x}, t) \quad \text{infinite set of generalized} \\ t &\rightsquigarrow t \quad \text{coordinates} \\ i &\rightsquigarrow i, \vec{x} \\ \sum_i &\rightsquigarrow \sum_i \int d^3\vec{x} \end{aligned}$$

Lagrangian  $L(q_i, \dot{q}_i, t) \rightsquigarrow$  Lagrangian density  $\mathcal{L}(\phi_i, \partial_\mu \phi_i, x^\mu)$

$L \equiv \int d^3x \mathcal{L}$  integrated over a spacelike time slice.

We assume that  $\mathcal{L}$  is local in space and time, and depends on at most first derivatives w.r.t.  $t, \vec{x}$ .

If  $\mathcal{L}$  is a Lorentz scalar, then the Euler-Lagrange eqs will be Lorentz covariant.

$$\delta S = \sum_i \int \underbrace{dt d^3x}_{d^4x} \left[ \frac{\partial \mathcal{L}}{\partial \phi_i} \delta \phi_i + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_i)} \delta (\partial_\mu \phi_i) \right]$$

$\equiv \pi_i^\mu \uparrow \quad \quad \quad \uparrow = \partial_\mu \delta \phi_i$

under arbitrary variations  $\delta \phi_i$  such that  $\delta \phi_i(t_1, \vec{x}) = \delta \phi_i(t_2, \vec{x}) = 0$ .

Integrate by parts  $\rightsquigarrow \delta S = \sum_i \int d^4x \left[ \frac{\partial \mathcal{L}}{\partial \phi_i} - \partial_\mu \pi_i^\mu \right] \delta \phi_i = 0$ .

$$\Rightarrow \frac{\partial \mathcal{L}}{\partial \phi_i} - \partial_\mu \pi_i^\mu = \frac{\partial \mathcal{L}}{\partial \phi_i} - \partial_\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_i)} = 0$$

Euler-Lagrange Eqs.

The canonical momentum is  $\pi_i(t, \vec{x}) \equiv \pi_i^0(t, \vec{x})$

(Don't think of  $\pi^{\mu}_i$  as a 4-vector generalization of the canonical momentum.)

$$\text{Hamiltonian } H = \int_i d^3x \left[ \pi_i \partial_0 \phi_i - \mathcal{L} \right]$$

$$\text{Hamiltonian density } \mathcal{H} = \sum_i \left[ \pi_i \partial_0 \phi_i - \mathcal{L} \right]$$

Example: Most general  $\mathcal{L}$  satisfying:

(1)  $\mathcal{L}$  is a Lorentz scalar.

(2)  $\mathcal{L}$  built from one real scalar field  $\phi = \phi^*$

(3)  $\mathcal{L}$  quadratic in  $\phi, \partial_\mu \phi$

↖ for linear eqs. of motion

$$\mathcal{L} = \frac{1}{2} a \left[ \partial_\mu \phi \partial^\mu \phi + b \phi^2 \right]$$

We can rescale  $\phi \rightarrow \frac{\phi}{\sqrt{|a|}}$ ,  $\mathcal{L} \rightarrow \pm \frac{1}{2} \left[ \partial_\mu \phi \partial^\mu \phi + b \phi^2 \right]$

$$\pi^{\mu} = \frac{\partial \mathcal{L}}{\partial(\partial_\mu \phi)} = \frac{\partial}{\partial(\partial_\mu \phi)} \left( \pm \frac{1}{2} \eta^{\alpha\beta} \partial_\alpha \phi \partial_\beta \phi \pm \frac{1}{2} b \phi^2 \right)$$

$$\text{Use } \frac{\partial(\partial_\alpha \phi)}{\partial(\partial_\mu \phi)} = \delta_\alpha^\mu$$

$$\begin{aligned} \pi^\mu &= \pm \frac{1}{2} \eta^{\alpha\beta} \left( \delta_\alpha^\mu (\partial_\beta \phi) + (\partial_\alpha \phi) \delta_\beta^\mu \right) \\ &= \pm \frac{1}{2} \left( \eta^{\mu\beta} (\partial_\beta \phi) + \eta^{\alpha\mu} (\partial_\alpha \phi) \right) \\ &= \pm \partial^\mu \phi \end{aligned}$$

$$\partial_\mu \pi^\mu - \frac{\partial \mathcal{L}}{\partial \phi} = 0 \Rightarrow \boxed{\pm (\partial_\mu \partial^\mu \phi - b \phi) = 0}$$

The Euler-Lagrange eqs. are the Klein-Gordon equations in this theory if we define  $b = -m^2$ .

⇒ Terms in  $\mathcal{L}$  quadratic in the fields without derivatives are related to masses.

$$\begin{aligned}\mathcal{H} = \pi^0 \partial_0 \phi - \mathcal{L} &= \pm \left[ \frac{1}{2} (\pi^0)^2 + \frac{1}{2} (\nabla \phi)^2 - \frac{1}{2} b \phi^2 \right] \\ &= \pm \frac{1}{2} \left[ (\pi^0)^2 + (\nabla \phi)^2 + m^2 \phi^2 \right]\end{aligned}$$

Each term in brackets is positive definite and can be made arbitrarily large. For stability we need  $m^2 > 0$ , and the overall positive sign.

$$\Rightarrow \boxed{\mathcal{L} = \frac{1}{2} \left( \partial_\mu \phi \partial^\mu \phi - m^2 \phi^2 \right)}$$

This Lagrangian density defines the free scalar field.

Plane wave solutions:  $\phi_{\vec{k}}(t, \vec{x}) = a_{\vec{k}} e^{-i(\omega t - \vec{E} \cdot \vec{x})} + c.c.$

$$\left( \partial_\mu \partial^\mu + m^2 \right) \phi_{\vec{k}}(t, \vec{x}) = 0 \leadsto \boxed{\omega^2 = \vec{k}^2 + m^2}$$

$m$  is called the mass of the free scalar field.

An arbitrary solution to the E-L eq. can be decomposed in plane waves.