

Problem Set 8: Warped Extra Dimensions - The Background

1. *Fat Branes*

Assume a 5D metric of the form  $ds^2 = e^{-A(y)} dx^\mu dx^\nu \eta_{\mu\nu} + dy^2$ . (As usual, Greek indices run from 0 to 3, and capital Latin indices will run from 0 to 4.)

- Calculate the Christoffel symbols  $\Gamma_{\mu\nu}^y$  and  $\Gamma_{\nu y}^\mu$ .
- Calculate the nonvanishing components of the Ricci tensor,  $R_{\mu\nu}$  and  $R_{yy}$ . *Hint:* The Ricci tensor can be written as,

$$R_{MN} = \partial_P \Gamma_{MN}^P - \partial_M \Gamma_{PN}^P + \Gamma_{MN}^Q \Gamma_{QP}^P - \Gamma_{PN}^Q \Gamma_{QM}^P .$$

- Calculate the curvature scalar,  $R = g^{MN} R_{MN}$ .
- Derive the form of the stress tensor  $T_{MN}$  which is consistent with the above metric and Einstein's equations,  $R_{MN} - \frac{1}{2} g_{MN} R = 8\pi G_N T_{MN}$ .
- Write the stress tensor in the form,

$$T_M^N = \begin{pmatrix} \Lambda(y) + V(y) & & & & \\ & \Lambda(y) + V(y) & & & \\ & & \Lambda(y) + V(y) & & \\ & & & \Lambda(y) + V(y) & \\ & & & & \Lambda(y) \end{pmatrix} .$$

What are  $\Lambda(y)$  and  $V(y)$ ? (*Hint:* Be careful. Remember,  $T_M^N = T_{MP} g^{PN}$ .)

- If  $A(y) = k|y|$  for some constant  $k$ , then derive the relation between the cosmological constant  $\Lambda$  and the brane tension  $V$  discussed in class. (*Hint:* Recall that  $8\pi G_N = 1/(2M_{Pl}^3)$  in 4+1 dimensions.)
- If  $A(y)$  is a smooth function that approximates  $k|y|$ , then plot (roughly)  $\Lambda(y)$  and  $V(y)$ .

You are welcome to use a symbolic manipulation program like Mathematica to help you, but it shouldn't be necessary.