

Fully quantized description of light-atom interaction

Reminder: semiclassical description (we treated e-m field as an external interaction)

$$\hat{H}_{sc} = \hat{H}_a + \hat{H}_{int} = \sum_i E_i |i\rangle \langle i| - \hat{\mathbf{d}} \cdot \vec{\mathbf{E}}$$

(atom)

using $\langle i | \hat{\mathbf{d}} \cdot \vec{\mathbf{e}}_p | j \rangle = p_{ij}$ where $\vec{\mathbf{e}}_p$ is the polarization vector

$$\hat{H}_{sc} = \sum_i E_i |i\rangle \langle i| - \sum_{\substack{i,j \\ i \neq j}} p_{ij} |i\rangle \langle j| \cdot \mathbf{E}$$

Quantized e-m field

Now the system consists of two quantum objects: an electron state of an atom and the state of the photons in e-m field.

$$\hat{H} = \hat{H}_a + \hat{H}_{ph} + \hat{H}_{int} \quad \hat{H}_{ph} = \hbar\omega \left(\hat{a}^\dagger \hat{a} + \frac{1}{2} \right)$$

$$\hat{H}_{int} = -\hat{\mathbf{d}} \cdot \vec{\mathbf{E}} \quad \vec{\mathbf{E}} = \sqrt{\frac{\hbar\omega}{2\epsilon_0 V}} \vec{\mathbf{e}}_p \left[\underbrace{\hat{a} e^{-i\omega t}}_{\hat{a}(t)} + \underbrace{\hat{a}^\dagger e^{i\omega t}}_{\hat{a}^\dagger(t)} \right]$$

$$\hat{H}_{int} = -\sqrt{\frac{\hbar\omega}{2\epsilon_0 V}} (\hat{\mathbf{d}} \cdot \vec{\mathbf{e}}_p) (\hat{a} + \hat{a}^\dagger)$$

The quantum state basis: $|i, n\rangle = |i\rangle |n\rangle$

$|i\rangle$ an atom is in state $|i\rangle$

$|n\rangle$ photons in the e-m field

In general $|\psi\rangle = \sum_i \sum_n C_{i,n} |i, n\rangle$

Let us now consider the interaction of a quantized e-m field with a two-level system ($|a\rangle, |b\rangle$)

If we assume that we start at the atomic state $|a\rangle$ having $|n\rangle$ photons

$$\begin{aligned} \hat{H}_{int} |a, n\rangle &= -\sqrt{\frac{\hbar\omega}{2\epsilon_0 V}} \rho_{ab} (\hat{a} + \hat{a}^\dagger) |a, n\rangle |b\rangle \langle a| \cdot |a\rangle \\ &= \left[-\sqrt{\frac{\hbar\omega}{2\epsilon_0 V}} \rho_{ab} \right] |b\rangle \left(\underbrace{\sqrt{n} |n-1\rangle}_{\text{photon absorption}} + \underbrace{\sqrt{n+1} |n+1\rangle}_{\text{photon emission}} \right) \end{aligned}$$

Thus, as time goes on, our system can find itself in three possible states

$$\begin{aligned} |\psi(t)\rangle &= c_a(t) |a\rangle |n\rangle e^{-iE_a t/\hbar} e^{-i\hbar\omega t} + \\ &+ c_b(t) |b\rangle |n-1\rangle e^{-iE_b t/\hbar} e^{-i(n-1)\omega t} + \\ &+ c'_b(t) |b\rangle |n+1\rangle e^{-iE_b t/\hbar} e^{-i(n+1)\omega t} \end{aligned}$$

(it is convenient to keep both absorption and emission here since we don't know if $E_a > E_b$ or $E_a < E_b$)

Perturbation calculations $\omega_{ba} = \frac{E_b - E_a}{\hbar}$

$$c_b^{(1)}(t) = -\frac{i}{\hbar} \int_0^t \langle b, n-1 | \hat{H}_{int} | a, n \rangle e^{+i(\omega_{ba} - \omega)t'} dt'$$

$$c_b'^{(1)}(t) = -\frac{i}{\hbar} \int_0^t \langle b, n+1 | \hat{H}_{int} | a, n \rangle e^{+i(\omega_{ba} + \omega)t'} dt'$$

$$\langle b, n \pm 1 | \hat{H}_{int} | a, n \rangle = -\sqrt{\frac{\hbar\omega}{2\epsilon_0 V}} \rho_{ba} \begin{cases} \sqrt{n+1} & "+" \\ \sqrt{n} & "-" \end{cases}$$

$$g_{ba} = \left[+\sqrt{\frac{\hbar\omega}{2\epsilon_0 V}} \right] \frac{\rho_{ba}}{\hbar} = \frac{E_0 \rho_{ba}}{\hbar}$$

single-photon Rabi frequency (or coupling constant)

Electric field of a single photon

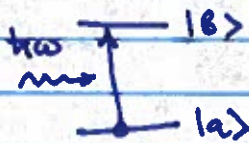
$$\langle b, n \pm 1 | \hat{H}_{int} | a, n \rangle = - \hbar g_{ba} \begin{cases} \sqrt{n+1} \\ \sqrt{n} \end{cases}$$

Total $|a\rangle \rightarrow |b\rangle$ transition amplitude

$$c_b(t) + c_b'(t) = g_{ba} \left\{ \sqrt{n} \frac{e^{i(\omega - \omega_{ba})t} - 1}{\omega - \omega_{ba}} + \sqrt{n+1} \frac{e^{i(\omega + \omega_{ba})t} - 1}{\omega + \omega_{ba}} \right\}$$

Near-resonant interaction $\omega \approx |\omega_{ba}| \rightarrow$ rotating wave approximation

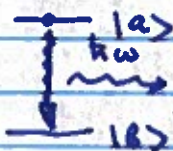
a) $E_b > E_a$ - the first term dominates



one photon is absorbed

$$P_{abs} = |c_b(t)|^2 \propto |g_{ba}|^2 \cdot n$$

b) $E_b < E_a$ - the second term dominates



one photon is emitted

$$P_{em} = |c_b'(t)|^2 \propto |g_{ba}|^2 \cdot (n+1)$$

$$\frac{P_{emission}}{P_{absorption}} = \frac{n+1}{n}$$

If $n=0$ (no light initially), the emission is still possible if $|a\rangle$ is an excited atomic state

\rightarrow spontaneous emission

Let's rewrite the interaction Hamiltonian in a form for a two-level system

$$\hat{H} = \sum_{i=a,b} E_i |i\rangle\langle i| + \hbar\omega \left(\hat{a}^\dagger + \hat{a} + \frac{1}{2} \right) + \hbar(g_{ba} |b\rangle\langle a| + g_{ab} |a\rangle\langle b|) \times (\hat{a} + \hat{a}^\dagger)$$

For the atomic component $\hat{H}_a = E_a |a\rangle\langle a| + E_b |b\rangle\langle b|$
 let's pick E_0 in the mid-point b/w $|a\rangle$ & $|b\rangle$

E_0 -----
 E_a -----
 $\omega_0 = \frac{E_b - E_a}{2}$

$$\hat{H}_a = -\frac{E_b - E_a}{2} |a\rangle\langle a| + \frac{E_b - E_a}{2} |b\rangle\langle b| =$$

$$= \frac{1}{2} \hbar\omega_0 (|b\rangle\langle b| - |a\rangle\langle a|) = \frac{1}{2} \hbar\omega_0 \hat{\sigma}_z$$

where $\hat{\sigma}_z = |b\rangle\langle b| - |a\rangle\langle a|$ inversion operator

$$\hat{H}_{int} = (\hbar g_{ab} |a\rangle\langle b| + \hbar g_{ba} |b\rangle\langle a|) (\hat{a} + \hat{a}^\dagger) = \hbar g (\hat{\sigma}_- + \hat{\sigma}_+) (\hat{a} + \hat{a}^\dagger)$$

where $\hat{\sigma}_+ = |a\rangle\langle b|$ and $\hat{\sigma}_- = |b\rangle\langle a|$
 are atomic transition operators

Three atomic operators $\hat{\sigma}_\pm$ and $\hat{\sigma}_z$ obey Pauli matrices commutation relationships

$$[\hat{\sigma}_+, \hat{\sigma}_-] = \hat{\sigma}_z \quad \& \quad [\hat{\sigma}_z, \hat{\sigma}_\pm] = 2\hat{\sigma}_\pm$$

The expectation values of these operators are related to atomic density matrix elements

$$\langle \psi | \hat{\sigma}_+ | \psi \rangle = \langle \psi | b \rangle \langle a | \psi \rangle = C_b^* C_a = \rho_{ab}$$

$$\langle \psi | \hat{\sigma}_z | \psi \rangle = \langle \psi | b \rangle \langle b | \psi \rangle - \langle \psi | a \rangle \langle a | \psi \rangle = \rho_{bb} - \rho_{aa}$$

Using this notation

$$\hat{H}_{int} = \hbar g (\hat{\sigma}_- + \hat{\sigma}_+) (\hat{a} + \hat{a}^\dagger)$$

$$\hat{H}_a = \frac{1}{2} \hbar\omega_0 \hat{\sigma}_z$$

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$$(\hat{\sigma}_+ + \hat{\sigma}_-)(\hat{a}^\dagger + \hat{a}) = \hat{\sigma}_+ \hat{a}^\dagger + \hat{\sigma}_+ \hat{a} + \hat{\sigma}_- \hat{a}^\dagger + \hat{\sigma}_- \hat{a}$$

$\begin{array}{cccc} \uparrow & \uparrow & \downarrow & \downarrow \\ \hline \hline \hline \hline \end{array}$

For $\omega \approx \omega_0$ there are only two processes with non-vanishing probability:

- ① a photon is absorbed and the atom is excited
- ② a photon is emitted and the atom is de-excited

RWA Hamiltonian

$$\hat{H} = \frac{1}{2} \hbar \omega_0 \hat{\sigma}_z + \hbar \omega \hat{a}^\dagger \hat{a} + \hbar g (\hat{\sigma}_- \hat{a}^\dagger + \hat{\sigma}_+ \hat{a})$$

If there is no interaction ($g=0$) \rightarrow atomic and photonic components are decoupled

Eigenstate $|a, n\rangle, |b, n\rangle$ for any n

We will assume a closed system:

- an atom can only be found in two states $|a\rangle$ & $|b\rangle$ $|a\rangle\langle a| + |b\rangle\langle b| = \hat{1}$
- the total energy of the system is conserved, so that the total number of excitations is constant

$$N_e = |b\rangle\langle b| + \hat{a}^\dagger \hat{a}$$

Two coupled states

$$|1\rangle = |b, n\rangle$$

$$E_{1,n}^{(0)} = \frac{1}{2} \hbar \omega_0 + \hbar \omega n = \hbar \omega (n + \frac{1}{2}) + \frac{1}{2} \hbar (\omega_0 - \omega)$$

$$|2\rangle = |a, n+1\rangle$$

$$E_{2,n}^{(0)} = -\frac{1}{2} \hbar \omega_0 + \hbar \omega (n+1) = \hbar \omega (n + \frac{1}{2}) - \frac{1}{2} \hbar (\omega_0 - \omega)$$

Energy splitting b/w $|1\rangle$ & $|2\rangle$ (in RWA)

$$\hbar \Delta = \hbar (\omega_0 - \omega)$$

So now we are back to a two-level system \rightarrow but for a fully quantized atom-photon states

for a known photon-number states

$$\hat{H}_n = \hbar\omega(n + \frac{1}{2}) \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} + \frac{1}{2}\hbar\Delta \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} + \hbar g \begin{pmatrix} 0 & \sqrt{n+1} \\ \sqrt{n+1} & 0 \end{pmatrix}$$

$$= \hbar\omega(n + \frac{1}{2}) \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} + \frac{1}{2}\hbar \begin{pmatrix} \Delta & 2g\sqrt{n+1} \\ 2g\sqrt{n+1} & -\Delta \end{pmatrix}$$

overall energy common for both states (not interesting) \rightarrow neglect

Quantum Rabi flopping

$$|\psi\rangle = c_a(t) |a, n+1\rangle + c_b(t) |b, n\rangle$$

we will assume $c_b(0) = 1, c_a(0) = 0$

$$i\hbar \frac{\partial \psi}{\partial t} = \hat{H} \psi$$

$$i\hbar \begin{pmatrix} \dot{c}_b \\ \dot{c}_a \end{pmatrix} = \frac{1}{2}\hbar \begin{pmatrix} \Delta & 2g\sqrt{n+1} \\ 2g\sqrt{n+1} & -\Delta \end{pmatrix} \begin{pmatrix} c_b \\ c_a \end{pmatrix}$$

~~this is~~ We have already encounter such problem for classical Rabi oscillations (with $2g\sqrt{n+1} \rightarrow \Omega$)

assuming $\Delta = 0$ for simplicity

$$c_b = \cos(g\sqrt{n+1}t)$$

$$c_a = -i \sin(g\sqrt{n+1}t)$$

$$|\psi(t)\rangle = \cos(g\sqrt{n+1}t) |b, n\rangle - i \sin(g\sqrt{n+1}t) |a, n+1\rangle$$

Quantum Rabi flopping

Note: even if $n=0$ (no photons), there still be oscillations @ frequency $g \rightarrow$ vacuum Rabi flopping

Thus, a Fock state with a fixed number of photons behaves very similar to a classical Rabi oscillations

What about a coherent state?

$$|\psi_{\text{atom}}\rangle|_{t=0} = |b\rangle, \quad |\psi_{\text{light}}\rangle|_{t=0} = \sum_{n=0}^{\infty} c_n |n\rangle = |d\rangle$$

$$c_n = e^{-|d|^2/2} \frac{|d|^n}{\sqrt{n!}}$$

$$|\psi_{\text{total}}\rangle|_{t=0} = |b\rangle|d\rangle$$

As we discussed before, the light atom interaction couples only pair of states $|b, n\rangle \leftrightarrow |a, n+1\rangle$, but now for all possible photon states $|n\rangle$

$$|\psi(t)\rangle = \sum_{n=0}^{\infty} \left(c_n \cos[g\sqrt{n+1}t] |b\rangle - i c_{n+1} \sin[g\sqrt{n+1}t] |a\rangle \right) |n\rangle$$

$$|\psi(t)\rangle = |\psi_a(t)\rangle + |\psi_b(t)\rangle$$

$$|\psi_a(t)\rangle = -i \sum_{n=0}^{\infty} c_{n+1} \sin(g\sqrt{n+1}t) |n+1\rangle$$

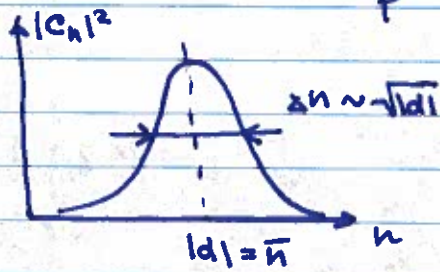
$$|\psi_b(t)\rangle = \sum_{n=0}^{\infty} c_n \cos(g\sqrt{n+1}t) |n\rangle$$

Average atomic inversion

$$\begin{aligned} \langle \psi(t) | \hat{\sigma}_z | \psi(t) \rangle &= \langle \psi(t) | (|b\rangle\langle b| - |a\rangle\langle a|) | \psi(t) \rangle = \\ &= \langle \psi_b | \psi_b \rangle - \langle \psi_a | \psi_a \rangle = \sum_{n=0}^{\infty} |c_n|^2 (\cos^2 g\sqrt{n+1}t - \sin^2 g\sqrt{n+1}t) \\ &= \sum_{n=0}^{\infty} |c_n|^2 \cos 2g\sqrt{n+1}t = e^{-|d|^2} \sum_{n=0}^{\infty} \frac{|d|^{2n}}{n!} \cos(2g\sqrt{n+1}t) \end{aligned}$$

The output is a combination of many oscillations of somewhat different periods \rightarrow no clear oscillations

Coherent state photon number distribution



Main contributions comes from components with frequencies b/w $2g\sqrt{n-\Delta n}$ and $2g\sqrt{n+\Delta n}$

Corresponding phase spread

$$2gt_c (\sqrt{n+\Delta n} - \sqrt{n-\Delta n}) \approx 2gt_c \sqrt{n} \left[\left(1 + \frac{\Delta n}{2n}\right) - \left(1 - \frac{\Delta n}{2n}\right) \right]$$

$$\approx 2gt_c \left(\frac{\Delta n}{\sqrt{n}} \right) \approx 2gt_c \Rightarrow t_c \sim 1/g$$

depends only on the coupling strength

However, we can also expect to see a revival of Rabi oscillations after some time t_R

$$(g\sqrt{n+1} - g\sqrt{n}) t_R = 2\pi \quad (\text{or in general } 2\pi k, k=1,2,\dots)$$

$$g\sqrt{n} \left(1 + \frac{1}{2n} - 1\right) t_R = \frac{g t_R}{2\sqrt{n}} = 2\pi$$

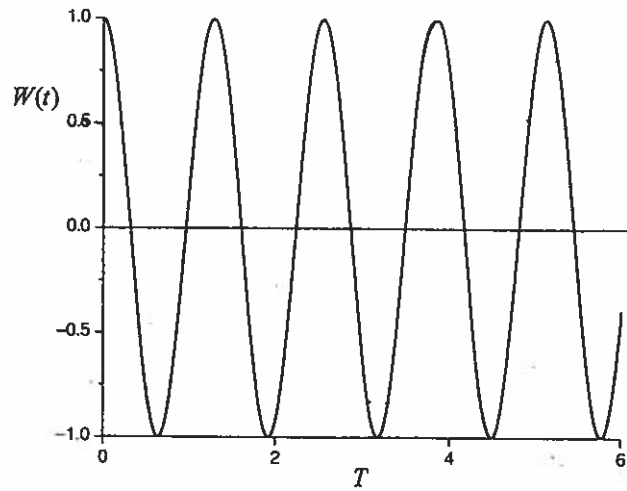
$$t_R \approx \frac{4\pi\sqrt{n}}{g}$$

The revival is never complete, since the frequencies $\{g\sqrt{n}\}$ are not truly equidistant?

Why a coherent state is less "classical" than a number state?

Clear Rabi flopping require knowledge of precise intensity, that is provided by the Fock states. A coherent state is a minimum uncertainty state, thus it has certain spread in its intensity distribution that leads to the Rabi flopping diffusion

Fig. 4.6. Periodic atomic inversion with the field initially in a number state $|n\rangle$ with $n = 5$ photons.



4.5 Fully quantum-mechanical model; the Jaynes-Cummings model

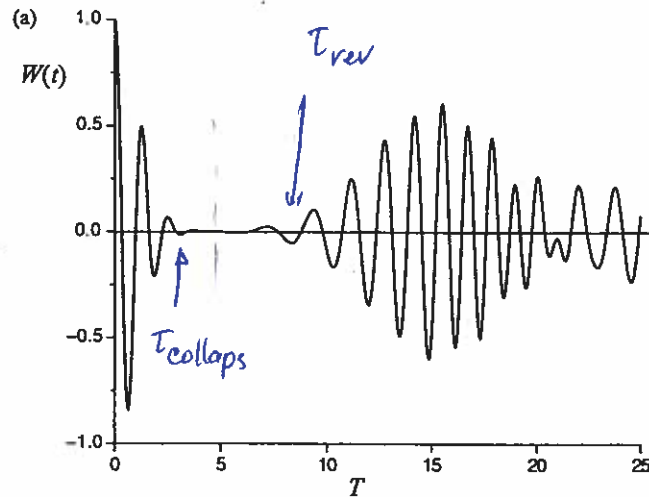


Fig. 4.7. (a) Atomic inversion with the field initially in a coherent state $\bar{n} = 5$. (b) Same as (a) but showing the evolution for a longer time, beyond the first revival. Here, T is the scaled time $\frac{t}{T}$.

