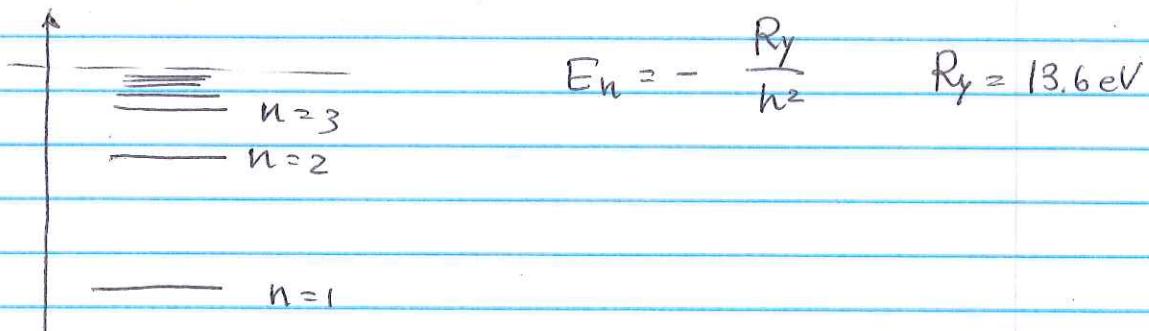


## Interaction of light with atoms: semiclassical description

semiclassical = quantized atoms + classical e-m field

From quantum ~~classical~~ quantum mechanics: discrete spectrum of electron states  
 Hydrogen atom (exactly solvable)  
 everything else - ap H-atom + corrections



Let us recall our classical oscillator model for a moment: atomic response  $\Rightarrow$  ~~refl~~  
 $m = \epsilon_0 \rho e^2 / \omega_0^2$  polarizability

$$m = \epsilon_0 \rho e^2 / \omega_0^2 \quad \text{polarizability}$$

and, for  $\delta \ll \omega_0$

$$\delta(\omega) \approx \frac{\epsilon_0 \rho e^2}{2 \epsilon_0 m \omega_0} \frac{1}{\omega_0 - \omega + i \delta/2}$$

Strong response only for  $\omega - \omega_0 \sim \delta$   
 if  $\omega$  is very different from  $\omega_0$   $d \rightarrow 0$

That behavior is also true for quantum system; the light interacts with electron stronger if its frequency is close to the energy difference  $\delta/\hbar$  between two transitions,  
 $\omega_0 \rightarrow (E_f - E_i) / \hbar$

That also means that even though there are many energy levels in atoms, only the two that are in resonance conditions ( $\omega \sim \omega_0$ ) are going to be important =>

Two-level approximation

Frequency of an atomic transition

$$--- 2$$

$$\omega_0 = (E_2 - E_1) / \hbar$$

Frequency of an e-m field

$\omega$ , close to  $\omega_0$

$$--- 1$$

Quantum state of atoms

$$|\psi\rangle = c_1 |1\rangle + c_2 |2\rangle$$

Probability ~~of~~ to find an atom in state 1

$$|c_1|^2, \text{ and in state 2} - |c_2|^2$$

If we have many identical atoms ( $N$ ), then  $N_1 = N \cdot |c_1|^2$  will be found in state 1, and  $N \cdot |c_2|^2$  - in state 2.

In a thermal equilibrium  $\frac{N_2}{N_1} \underset{R_B}{\approx} N_1 \sim e^{-E_1/k_B T}$

$$\text{and } \frac{N_2}{N_1} = e^{-(E_2-E_1)/k_B T}$$

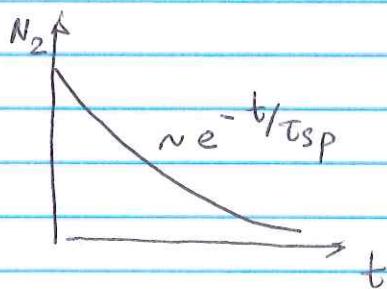
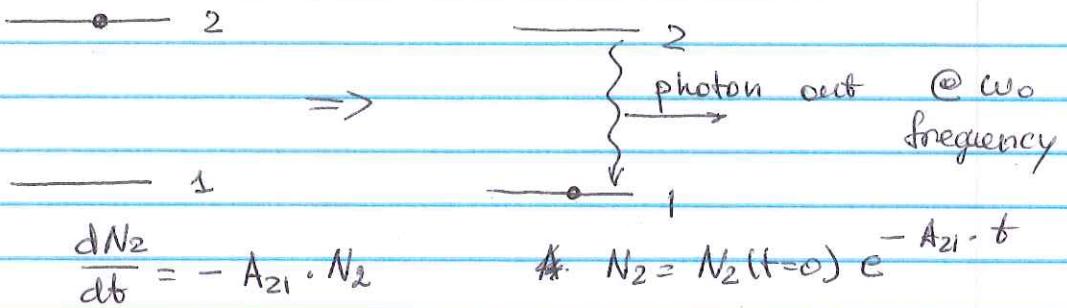
typically most atoms are in lower-energy state. In fact, in optical transitions

the higher states are basically empty

$$k_B T \approx 0.03 \text{ eV} \quad \text{and} \quad E_2 - E_1 \sim 1 \text{ eV} \gg k_B T$$

### Three important light - atom interactions

#### Spontaneous emission

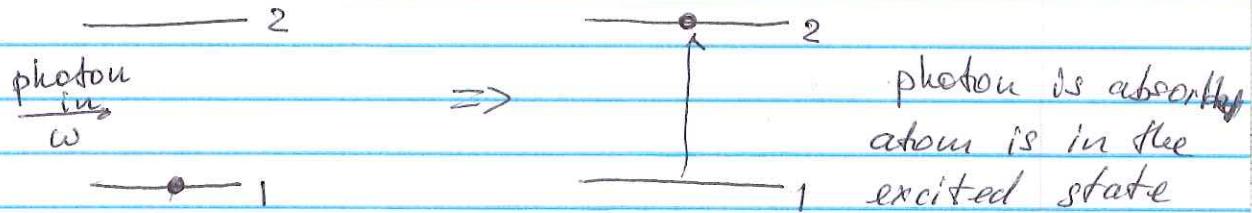


Spontaneous lifetime  
of the excited state

$$N_2 = N_2(t=0) e^{-t/\tau_{sp}}$$

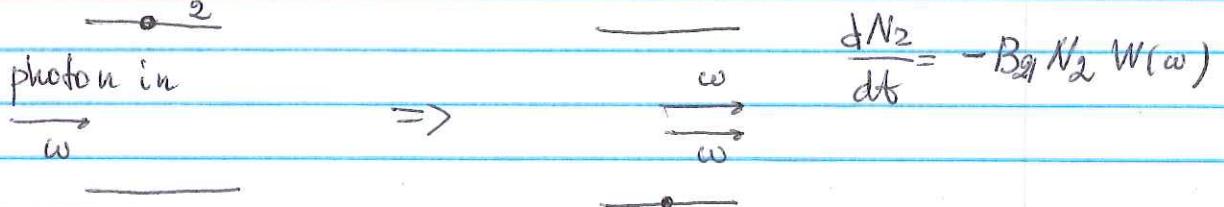
$$\tau_{sp} = 1/A_{21}$$

#### Stimulated absorption



$$\frac{dN_2}{dt} = B_{12} N_1 \cdot \underbrace{W(\omega)}_{\text{spectral energy density}} \quad (\text{for absorption})$$

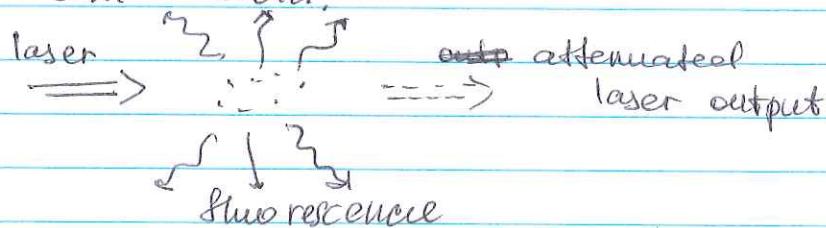
#### Stimulated emission



Regular atoms: all atomic population at the ground state, no atoms in the excited state  
 - absorption is the strongest

(starts populating state 2)

But if the field is weak, so that stimulated emission is not strong enough, the spontaneous emission overtakes, ~~the~~ and energy is dissipated via spontaneous emission from the incoming e-m field.



$A_{21}$ ,  $B_{12}$  and  $B_{21}$  - Einstein coefficient

$$B_{12} = B_{21}, \text{ and } \frac{A_{21}}{B_{21}} = \frac{8\pi\hbar f^3}{c^3} \quad f = \frac{c_0}{2\pi}$$

$1/A_{21} = \tau_{sp}$   $\rightarrow$  lifetime of an excited quantum state, can be calculated using advanced quantum mechanics.

Can we amplify the incoming e-m wave?

Yes! If stimulated emission is the dominant process

Assuming spontaneous emission negligible for e-m field travelling from  $z$  to  $z+dz$

$$\begin{aligned} \text{photon loss: } & dN_{\text{photons}}^{(-)} = -dN_2 = -B_{12} N_1 W(\omega) dz/c \\ \text{photon gain: } & dN_{\text{photons}}^{(+)} = -dN_2 = B_{21} N_2 W(\omega) dz/c \end{aligned}$$

Each photons carries energy  $h\nu$

Total energy balance

$$dW = \hbar\omega (B_{21}N_2 W - B_{12}N_1 W) \cdot \frac{dz}{c} =$$

$$= \frac{\hbar\omega}{c} B_{21} (N_2 - N_1) W dz$$

$$\frac{dW}{dz} = \frac{\hbar\omega}{c} B_{21} (N_2 - N_1) W$$

$$W(z) = W(z=0) e^{g z} \quad g - \text{gain}$$

$$g(f) = A_{21} (N_2 - N_1) \frac{c^2}{2\pi f^2} \quad (F(f)) \quad \text{Spectral line of the absorption}$$

To achieve gain need  $N_2 > N_1$   
population inversion

LASER - light amplification by the  
Stimulated Emission of Radiation