

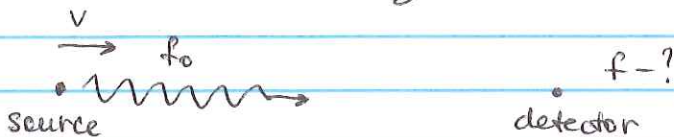
# Relativistic Doppler effect

(describes frequency change of an EM wave if a source (or detector) is moving)

Brief reminder: light is an electro-magnetic wave

$$E = E_0 \cos\left(\frac{2\pi}{\lambda} \cdot x - 2\pi f \cdot t\right) \quad (\text{in vacuum})$$

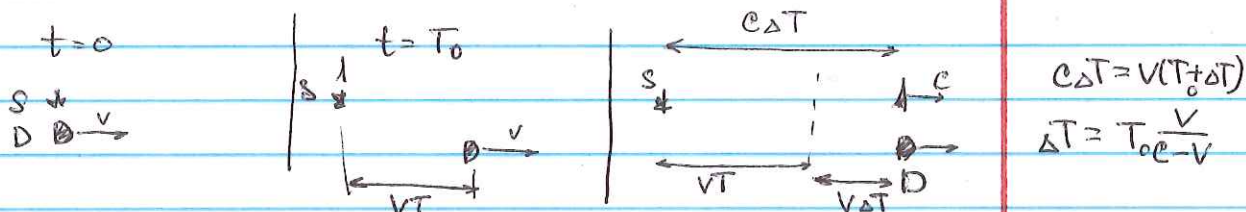
$f$  - frequency,  $\lambda$  - wavelength,  $f = c/\lambda$



Light clocks: let's assume that every time the light pulse hits the bottom mirror, a flash of light is emitted, to mark each "tick" (duration  $T_0$ )  
If a detector moves away from the source (clock) how much time will pass between each tick?



Let's for simplicity assume that at  $t=0$  the source and the detector are collocated



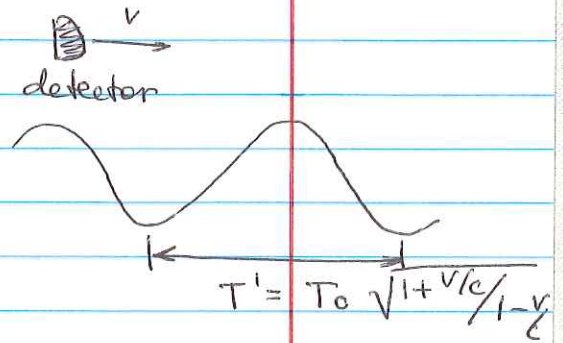
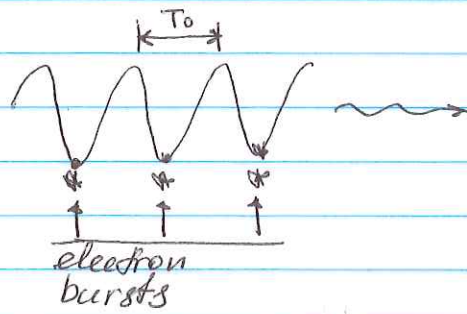
In the RF of the source the time b/w two detected ticks is  $T = T_0 \frac{v}{c-v} + T_0 = T_0 \frac{c}{c-v}$

In the RF of the moving detector, however, the time is dilated, so it perceives the ticks coming in time intervals

$$T' = \frac{T}{\gamma} = \frac{1}{\gamma} T_0 \frac{c}{c-v} = T_0 \sqrt{(1 - v^2/c^2)} \frac{1}{1 - v/c} = T_0 \sqrt{\frac{1 + v/c}{1 - v/c}}$$

Detector sees ticks coming closer to each other if the detector approaches the source (or vice versa), and they become ~~more~~ longer if the detector is receding

We can use the same logic to deduce the relativistic changes in EM frequency



Frequency of the source  $f_0 = \frac{1}{T_0}$

Frequency ~~of the~~ measured by the detector

$$f = \frac{1}{T^1} = \frac{1}{T_0} \sqrt{\frac{1-v/c}{1+v/c}} = f_0 \sqrt{\frac{1-v/c}{1+v/c}}$$

If the source and the detector are moving away  $f = f_0 \sqrt{\frac{1-v/c}{1+v/c}} < f_0$   $\lambda_0 = \frac{c}{f_0}$ ;  $\lambda = \lambda_0 \sqrt{\frac{1+v/c}{1-v/c}} > \lambda_0$

The detected light is red-shifted

If the detector is approaching  $v \rightarrow -v$

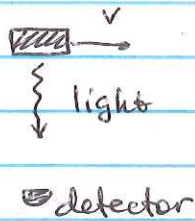
$$f = f_0 \sqrt{\frac{1+v/c}{1-v/c}} > f_0 \text{ or } \lambda = \lambda_0 \sqrt{\frac{1-v/c}{1+v/c}} < \lambda_0$$

The detected light is blue-shifted

(Joke: red traffic will appear green if you approach it fast enough)

Since the universe is expanding, distant galaxies move away from us, so their light is red-shifted.

## Transverse Doppler effect



No correction for the source motion, still time dilatation

$$f'_{\text{transverse}} = \frac{1}{\gamma} f_0 = \sqrt{1 - v^2/c^2} \cdot f_0$$

How big is the Doppler effect if the source/detector relative velocity is not close to  $c$ ?

If  $v/c \ll 1$ , we can use Taylor expansion to find the first order correction to the frequency

$$(1+x)^d \approx 1 + dx + \frac{d(d-1)}{2!} x^2 + \dots \quad \text{if } x \ll 1$$

if  $d = \pm 1/2$

$$(1 \pm x)^{1/2} \approx 1 \pm \frac{x}{2} \quad (1 \pm x)^{-1/2} \approx 1 \mp \frac{x}{2}$$

$$\sqrt{\frac{1-v/c}{1+v/c}} \approx \left(1 - \frac{1}{2} \frac{v}{c}\right) \left(1 + \frac{1}{2} \frac{v}{c}\right) \approx 1 - \frac{v}{c} \quad (\text{or } 1 + \frac{v}{c} \text{ for } v \rightarrow -v)$$

$$f = f_0 \sqrt{\frac{1-v/c}{1+v/c}} \approx f_0 - f_0 \frac{v}{c}$$

$$f = f_0 \sqrt{\frac{1+v/c}{1-v/c}} \approx f_0 + f_0 \frac{v}{c}$$

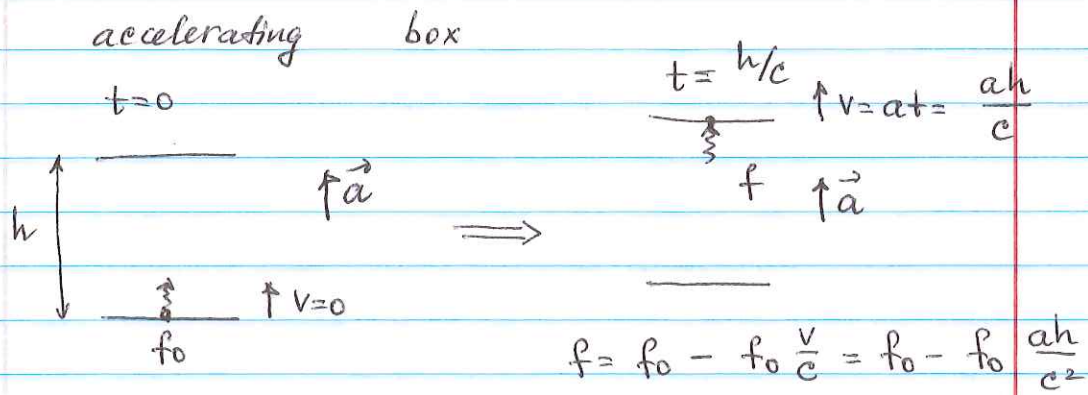
Doppler shift  
positive for approaching  
sources, negative for  
receding.

But for the transverse Doppler effect

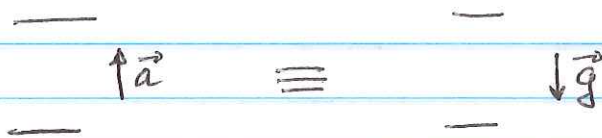
$$f'_{\text{transv}} = \frac{f_0}{\gamma} = \sqrt{1 - v^2/c^2} f_0 \approx f_0 - \frac{v^2}{2c^2} f_0$$

by a factor of  $\frac{v^2}{c^2} \ll 1$ .

# Gravitational red shift



Equivalence Principle: there is no experiment done in a small confined space which can tell the difference between a uniform gravitational field and an equivalent acceleration



Gravitational red shift  $\Delta f = - f_0 \frac{gh}{c^2}$